

Physics 307L

Spring 2021

Prof. Darcy Barron

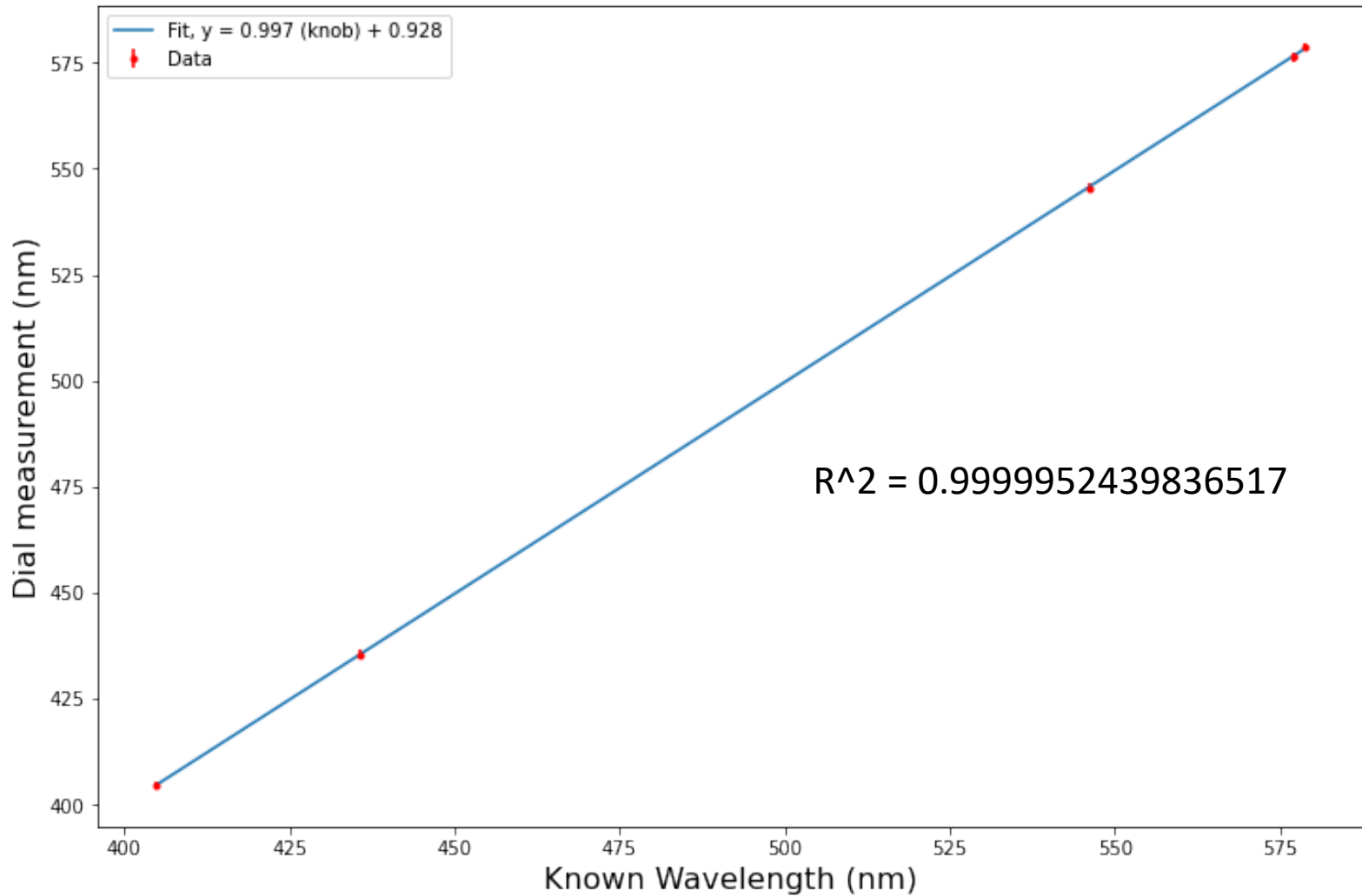
This lecture will be recorded

Updated Schedule

- Schedule of assignment due dates through the end of the semester is now posted on wiki, and will be in Teams soon
 - https://ghz.unm.edu/juniorlab/index.php?title=Schedule_Spring_2021#Course_Schedule
- **First lab report due date is this Wed, March 31**
 - Please ask if you have questions about completing the analysis for the experiment you write about in Lab Report 1
 - Lab report cannot be on Lab 0
 - https://ghz.unm.edu/education/juniorlab_pdfs/labreportguidelines.pdf
- **Schedule for Talk 2 is posted**
- **All students will give Talk 3** during scheduled during final exam time for this class, Friday, May 14 from 12:30pm – 2:30pm
 - This should not conflict with any other scheduled finals, but please let me know if there is a problem

Balmer Series Analysis

- Straightforward steps
 - Estimating uncertainties from equipment
 - Repeating measurements to estimate uncertainty
 - Propagating uncertainties
 - Simple linear fit to data
- More complex steps
 - Rejecting data (Chapter 6 of Taylor)
 - Choosing how to combine separate measurements
 - Least-squares fitting with errors in both dimensions



The slope = 0.9974273274123038, with uncertainty 0.0012558490265354887 The intercept = 0.9282554690792497, with uncertainty 0.6452914621011211

Chauvenet's Criterion

- If you make N measurements of a single quantity x , Chauvenet's criterion gives a simple test for deciding whether to reject a 'suspect value'
- $t_{sus} = \frac{|x_{sus} - \bar{x}|}{\sigma_x}$
- $n = N \times Prob(\textit{outside } t_{sus}\sigma)$
 - Use Appendix A to look up values
- If $n < 0.5$, then it is reasonable to reject x_{sus}

Example

- We make 10 measurements of length, x , and get these results:
 - 46, 48, 44, 38, 45, 47, 58, 44, 45, 43
 - What is the mean of the dataset? **45.8**
 - What is the standard deviation? **5.1**
 - What is the suspicious value? **58**
 - What is probability that such an outlier would appear from random chance?
 - **0.016**

Example 2

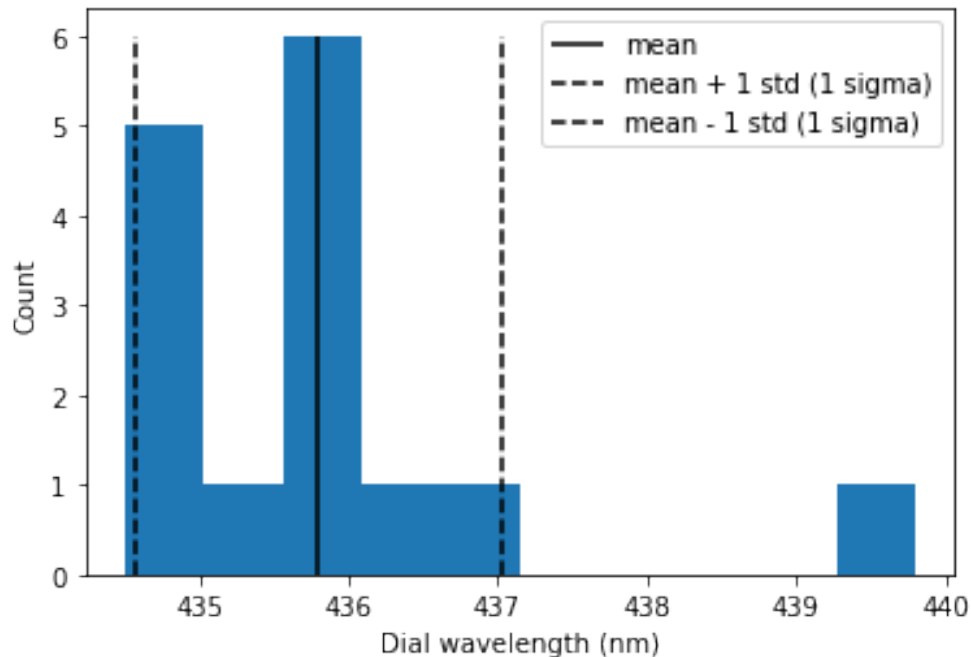
- Multiple measurements of second shortest wavelength line of mercury emission
- [437, 435.0, 435, 434.5, 439.8, 435.4, 435.6, 435.7, 436.2, 435.7, 436.0, 435.7, 435.8, 434.8, 434.8]
- What is mean? What is standard deviation?
 - **435.8, 1.23**
- What value is suspect? **439.8**
- Should it be thrown out and why?

Example 2

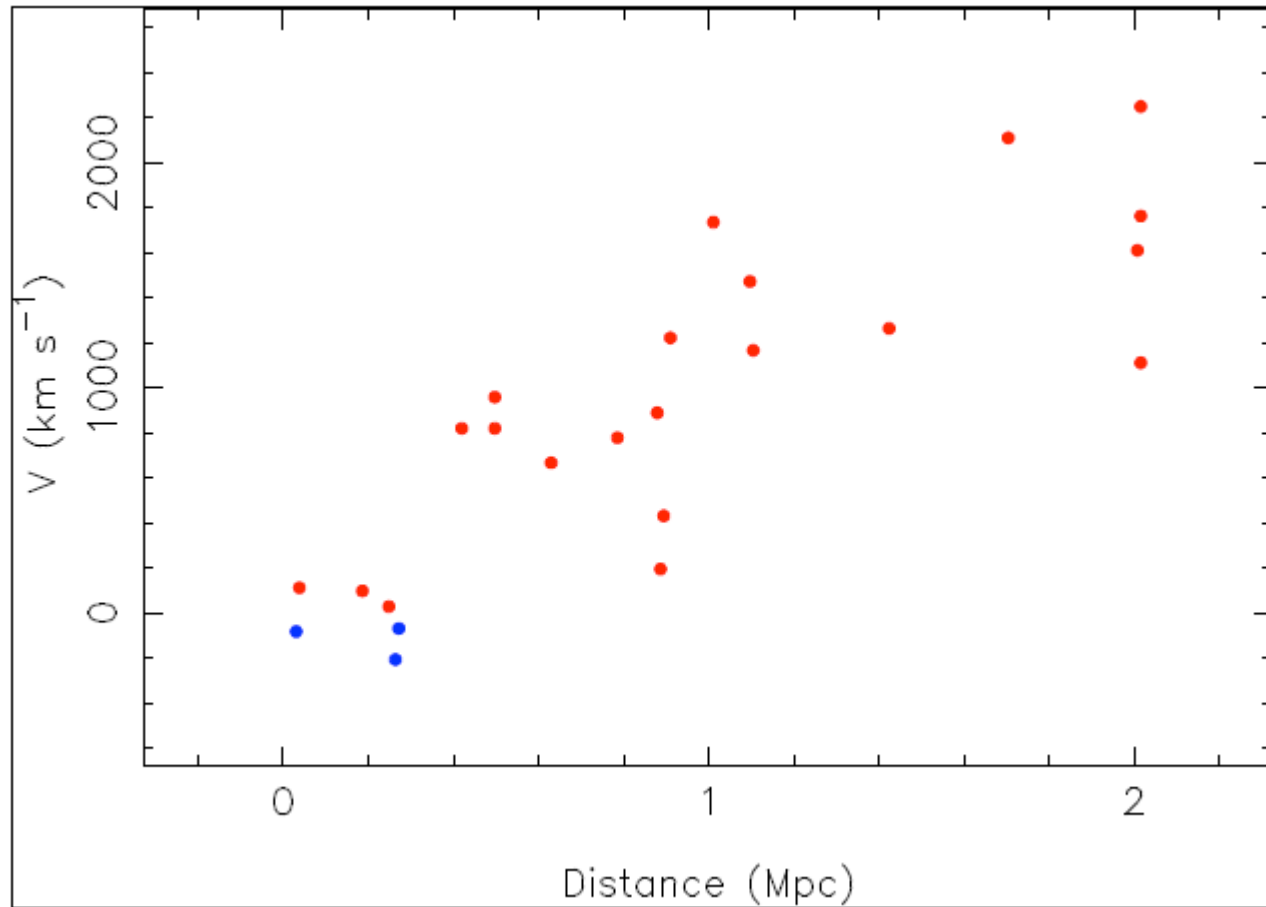
- Multiple measurements of second shortest wavelength line of mercury emission
- [437, 435.0, 435, 434.5, 439.8, 435.4, 435.6, 435.7, 436.2, 435.7, 436.0, 435.7, 435.8, 434.8, 434.8]
- Should it be thrown out and why?
 - Probably!
 - t-score is 3.24, so this is a 3-sigma outlier
 - 99.73% of random fluctuations will fall within 3 sigma of mean
 - So, there is a less than 0.27% chance one value would fluctuate this far away
 - We have 15 values, so there's a 4% chance it could have happened in this dataset
 - There is reasonable suspicion that the value might have been written down incorrectly or some other fluke, so it seems ok to throw it out

Example 2

- Multiple measurements of second shortest wavelength line of mercury emission
- [437, 435.0, 435, 434.5, 439.8, 435.4, 435.6, 435.7, 436.2, 435.7, 436.0, 435.7, 435.8, 434.8, 434.8]



Fitting data to a model



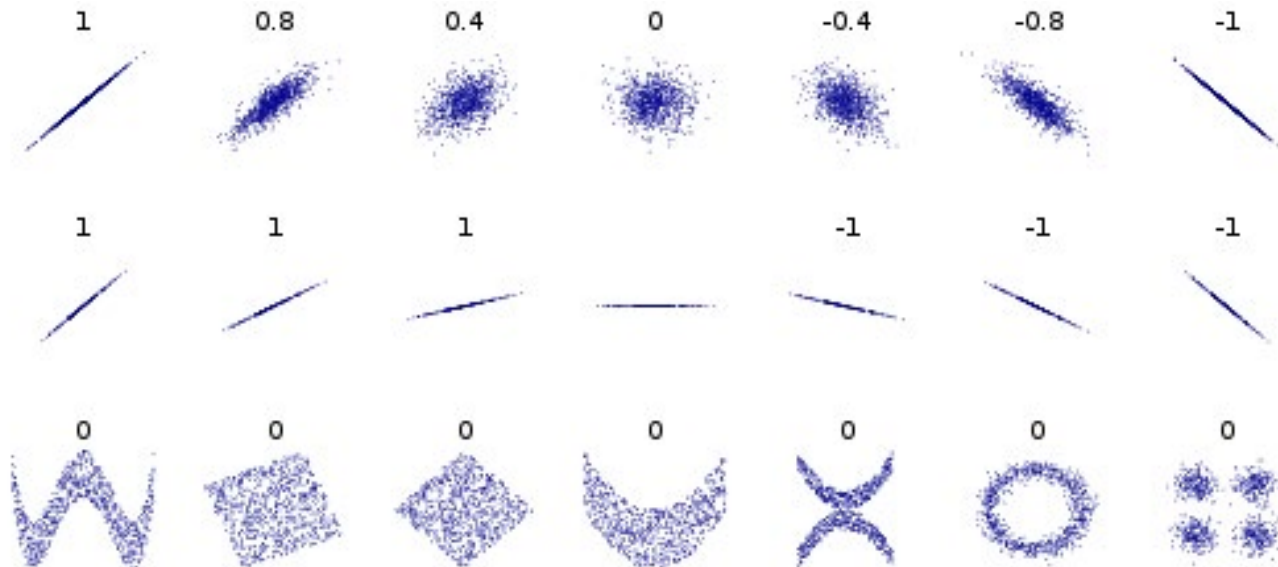
An early Hubble diagram, N=24 galaxies (1936)

Simplifying correlations

- **Linear correlations are easy to plot and examine**
- Linear fit
 - $y = mx + b$
- Can linearize your data to make it a linear correlation
 - Example: Balmer Series
 - $1/\lambda = R [1/2^2 - 1/n^2]$
 - Set $y = 1/\lambda$
- Proportional correlation is special case of linear fit where y-intercept is fixed at zero
 - $y = mx$

Pearson's correlation coefficient

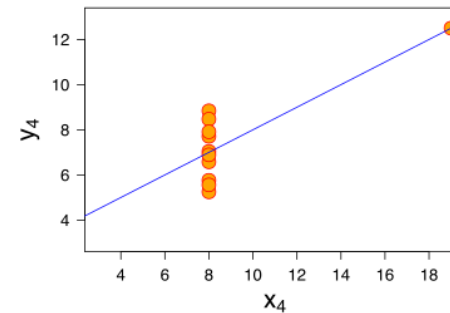
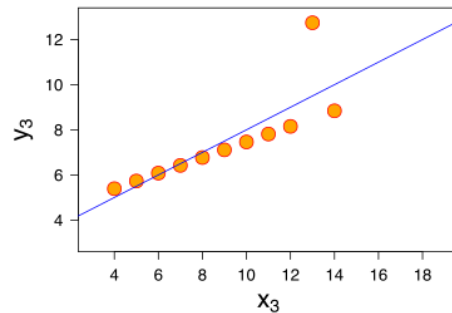
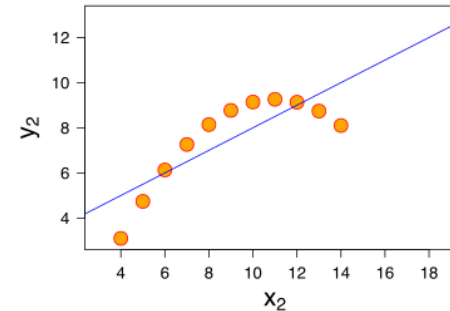
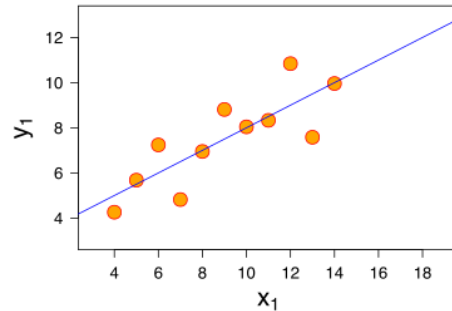
https://en.wikipedia.org/wiki/Pearson_correlation_coefficient



- Pearson's correlation coefficient, r , measures the linear correlation between two variables

Linear correlations are easy to plot and examine

Anscombe's quartet



Property	Value	Accuracy
Mean of x	9	exact
Sample variance of x	11	exact
Mean of y	7.50	to 2 decimal places
Sample variance of y	4.125	± 0.003
Correlation between x and y	0.816	to 3 decimal places
Linear regression line	$y = 3.00 + 0.500x$	to 2 and 3 decimal places, respectively
Coefficient of determination of the linear regression	0.67	to 2 decimal places

Plotting data can provide insight on trends not captured by statistics

https://en.wikipedia.org/wiki/Anscombe%27s_quartet

Speed of Light Analysis

- For our setup measuring the speed of light, we strongly expect to get data that fits a straight line
- For most ways of recording data, the data follows a linear fit with $y = mx + b$
- For some ways of recording data, it is a proportional fit $y = mx$
- What is the meaning of the slope?
- What is the meaning of the y-intercept?

Speed of light experimental design

- Without recording any data, we can say a few things about the capabilities and limitations of the experiment
 - We are performing the experiment on an optics table, so our maximum length is of order 1 meter, and our uncertainty is of order 0.5 cm
 - Placement error gets multiplied by 2 because roundtrip
 - We are using a fast pulsed laser and a fast oscilloscope, which sets our time resolution to ~ 0.5 ns
 - Are we limited by the laser, or the oscilloscope, or...?
 - How well calibrated is the oscilloscope?
- Expect ~ 5 -10% in systematic errors in our measurement of the speed of light in one leg of the optical path, limited by uncertainties in time measurement
- Can measure time delay added by adding something to optical path (acrylic rod), and derive index of refraction (with length)

Least-squares fitting

- Procedure described in Taylor Chapter 8
- Fitting multiple data points to a line can tell us how well the data is described by a linear fit, and gives a measure of statistical (random) fluctuations in our data set, or any lurking remaining systematic errors
- Just like with statistics (mean, standard deviation), knowing how it is done is important, so good to do it 'by hand' a few times
- However, also important to learn (*and understand*) at least one way to do it through programming
 - Excel, Python, Matlab, etc.
 - <https://ghz.unm.edu/juniorlab/index.php?title=ErrorAnalysis>
 - Can get into many variations once you start to weight data (touched on in Chapter 7 – weighted averages)

Example in Taylor

Section 8.2 Calculation of the Constants A and B

Table 8.1. Masses m_i (in kg) and lengths l_i (in cm) for a spring balance. The “ x ” and “ y ” in quotes indicate which variables play the roles of x and y in this example.

Trial number i	“ x ” Load, m_i	“ y ” Length, l_i	m_i^2	$m_i l_i$
1	2	42.0	4	84
2	4	48.4	16	194
3	6	51.3	36	308
4	8	56.3	64	450
5	10	58.6	100	586
$N = 5$	$\Sigma m_i = 30$	$\Sigma l_i = 256.6$	$\Sigma m_i^2 = 220$	$\Sigma m_i l_i = 1,622$

Example in Taylor

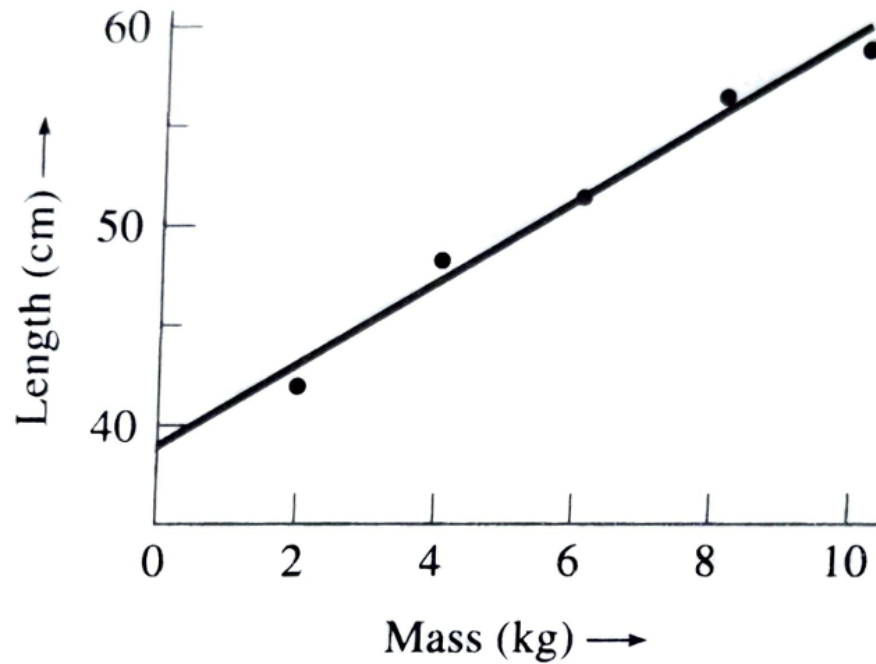
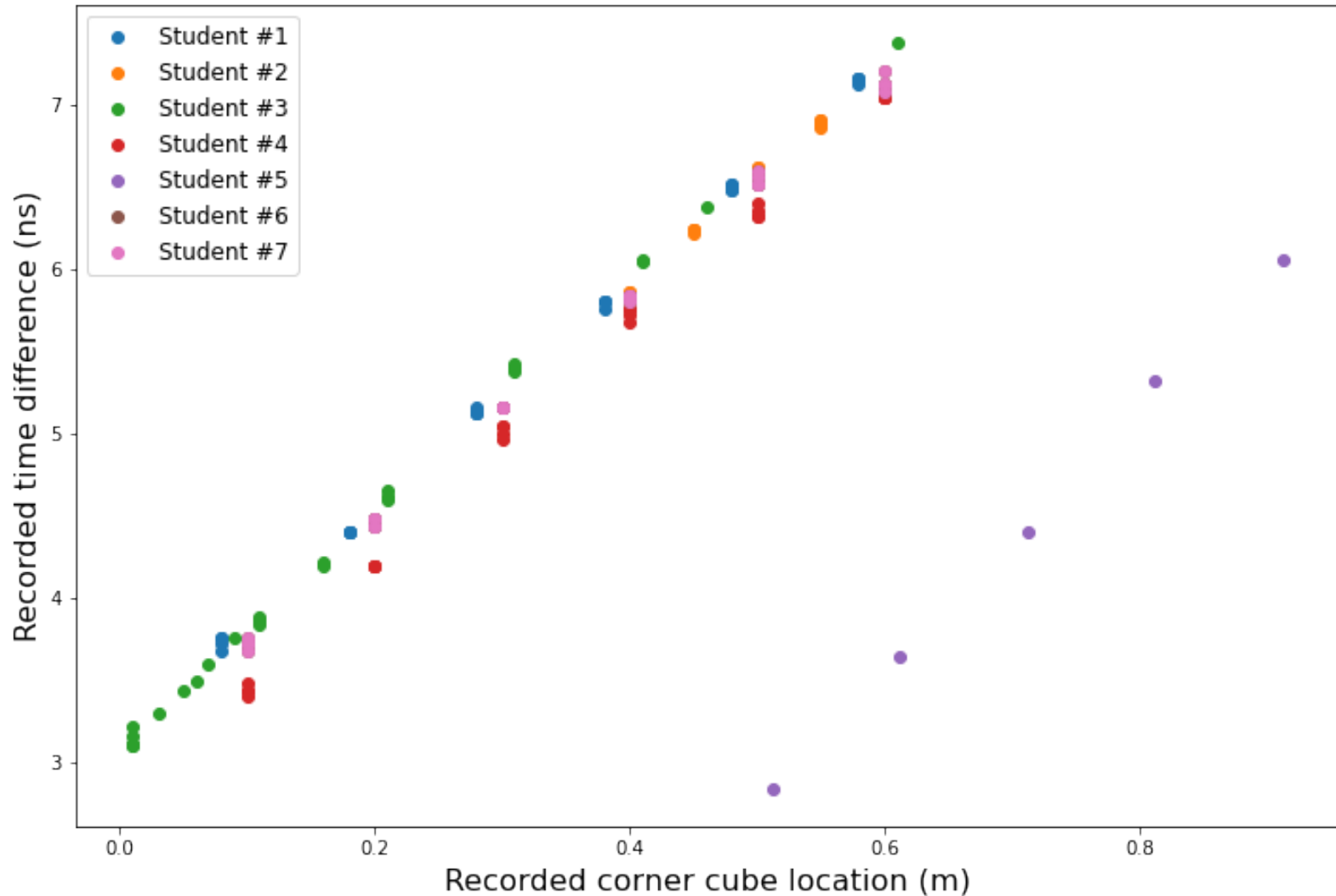
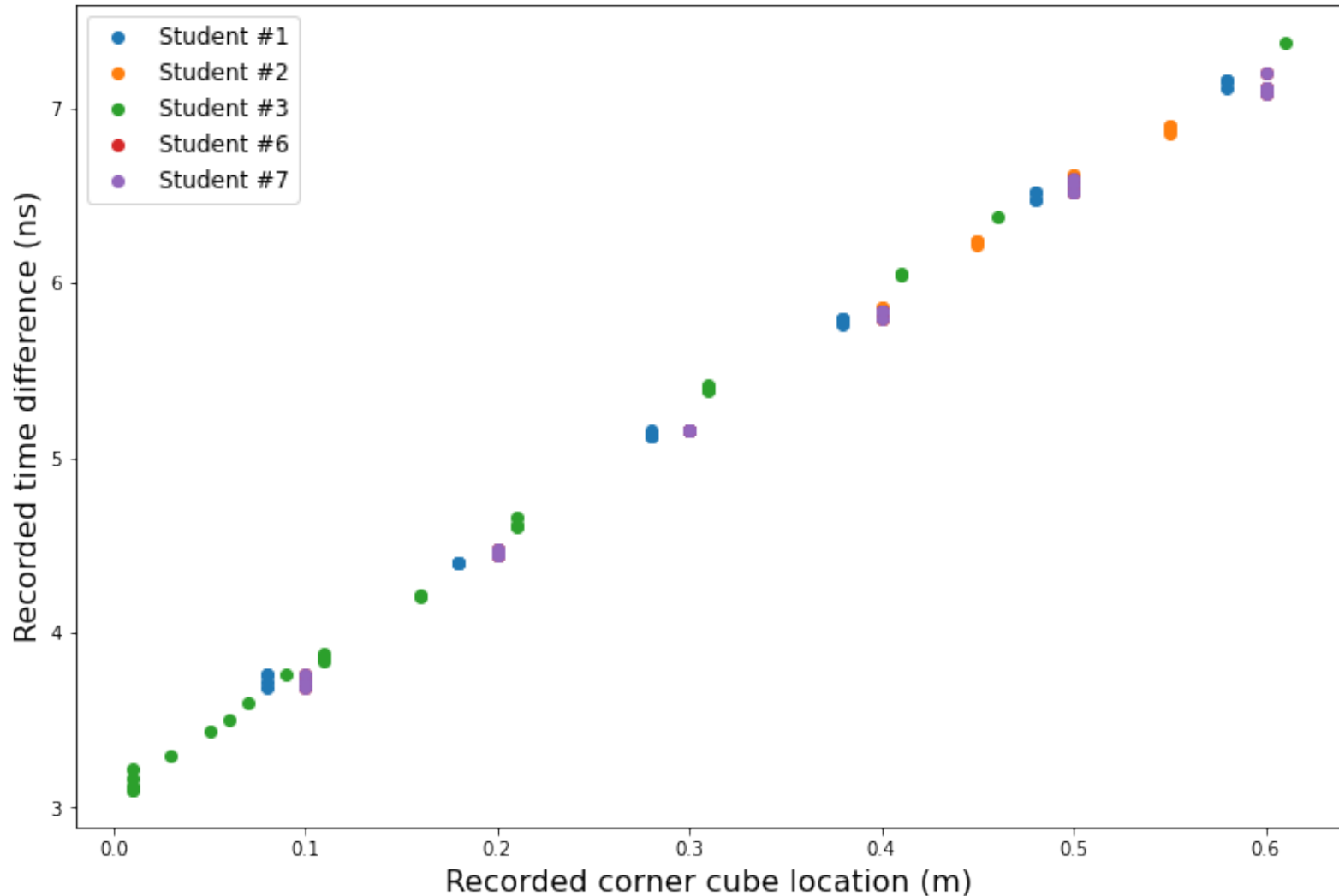


Figure 8.2. A plot of the data from Table 8.1 and the best-fit line (8.13).

Speed of Light Data – Our Class

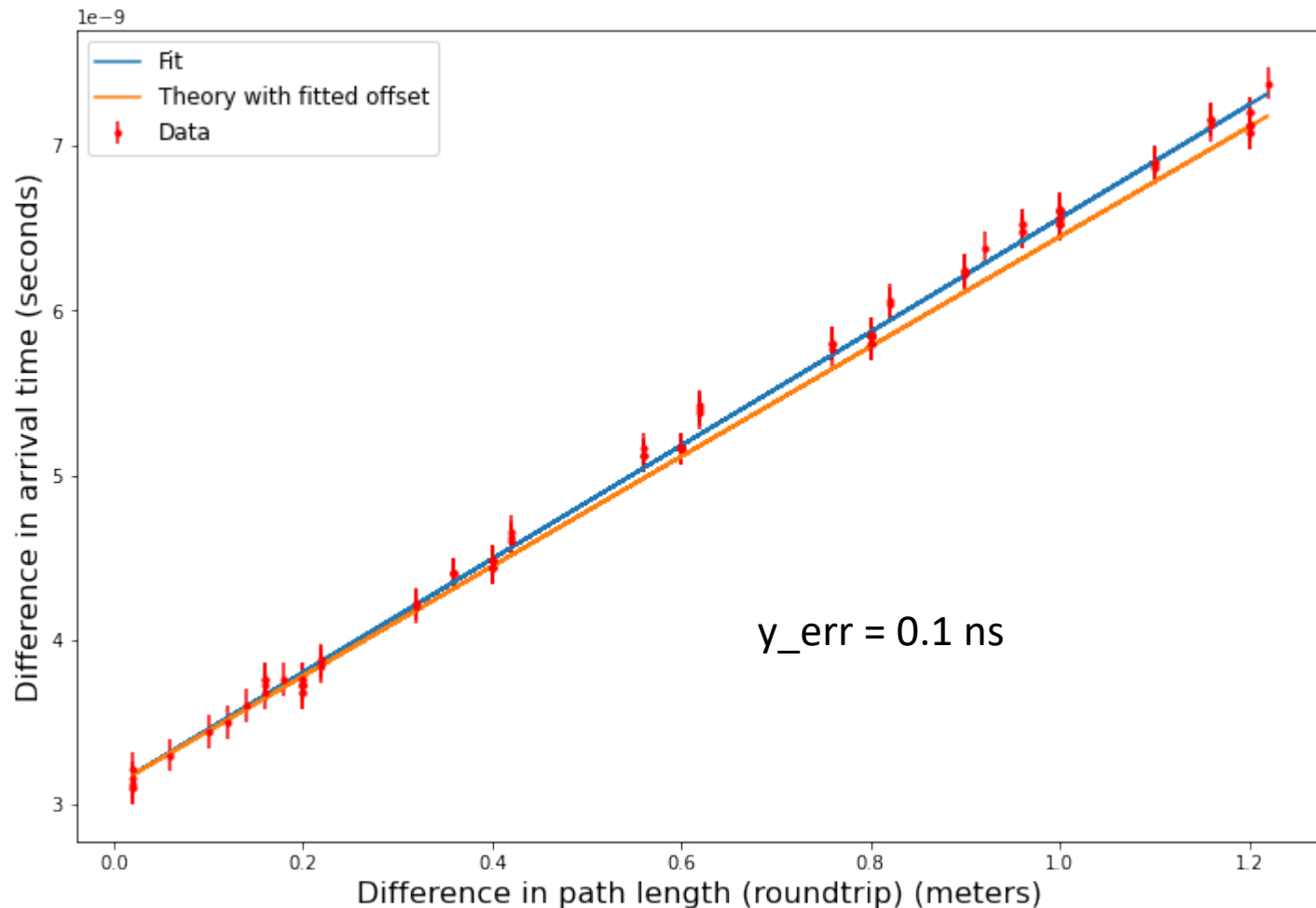


Speed of Light Data – Our Class Scrubbed of Inconsistent Data



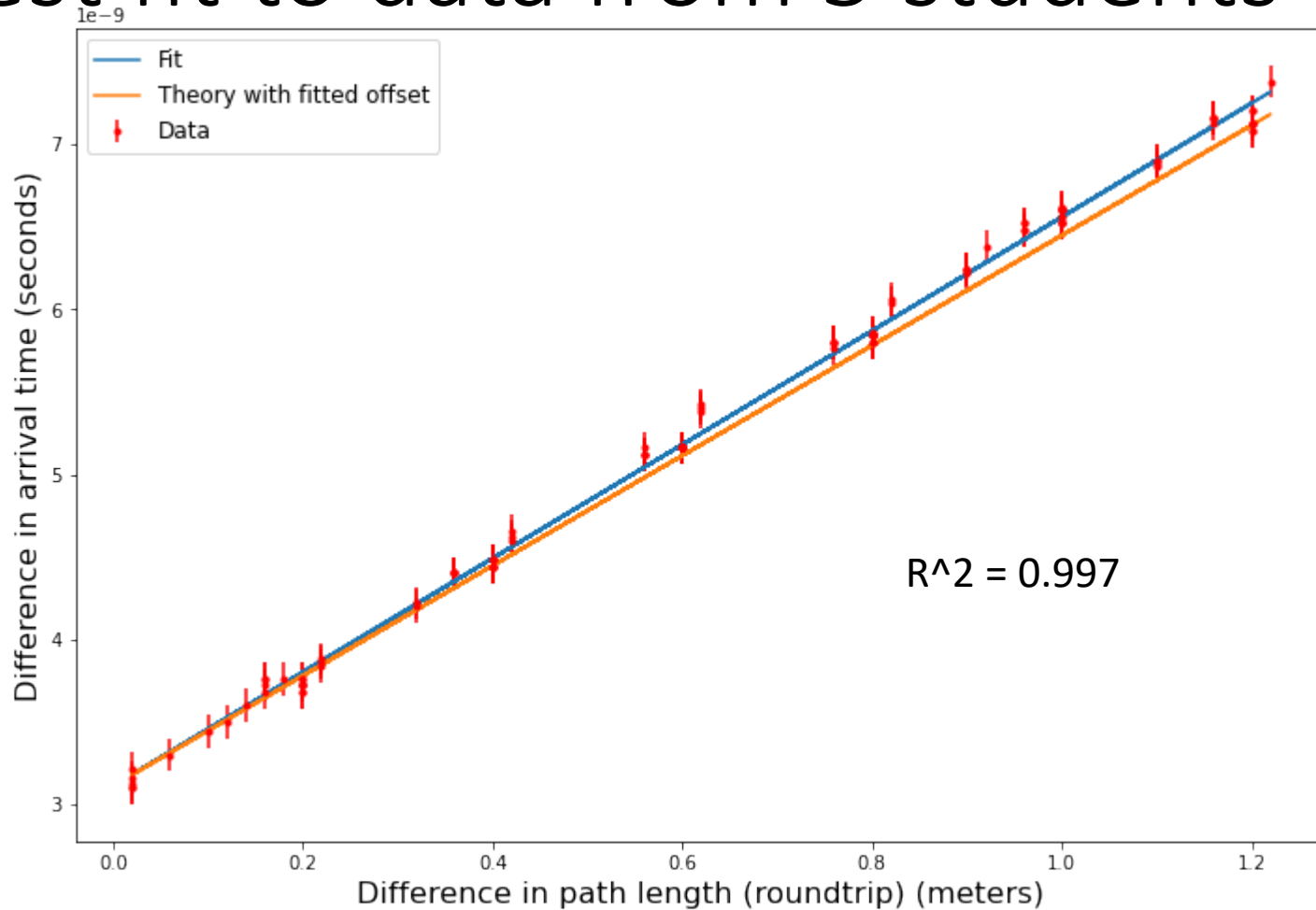
Speed of Light Data – Our Class

Best fit to data from 5 students



Speed of Light Data – Our Class

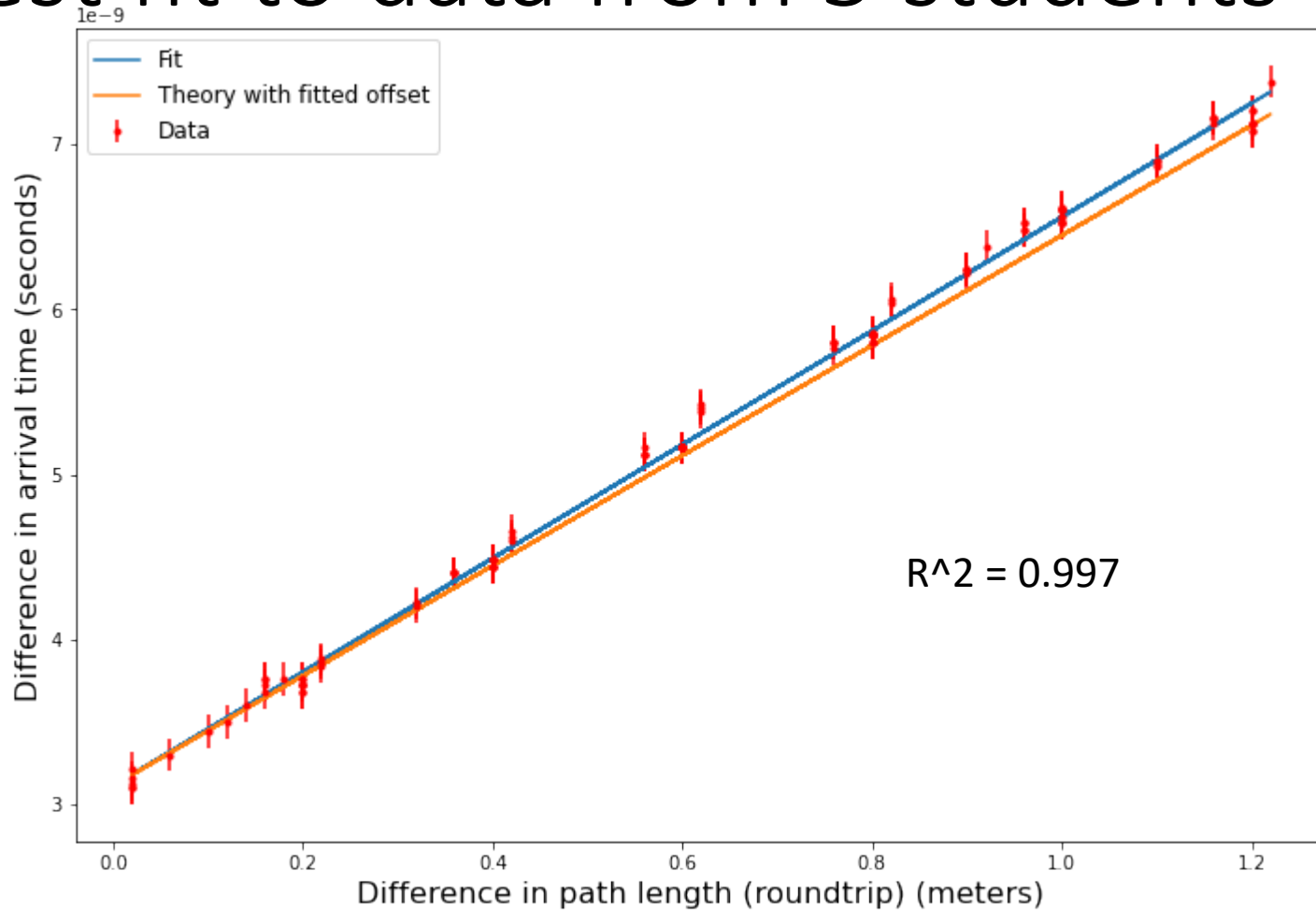
Best fit to data from 5 students



The slope = $3.4470692242258648e-09$, with uncertainty $1.741025266191759e-11$
The intercept = $3.111703798698238e-09$, with uncertainty $1.298000077865901e-11$
The speed of light = 290101513.76480633

Speed of Light Data – Our Class

Best fit to data from 5 students

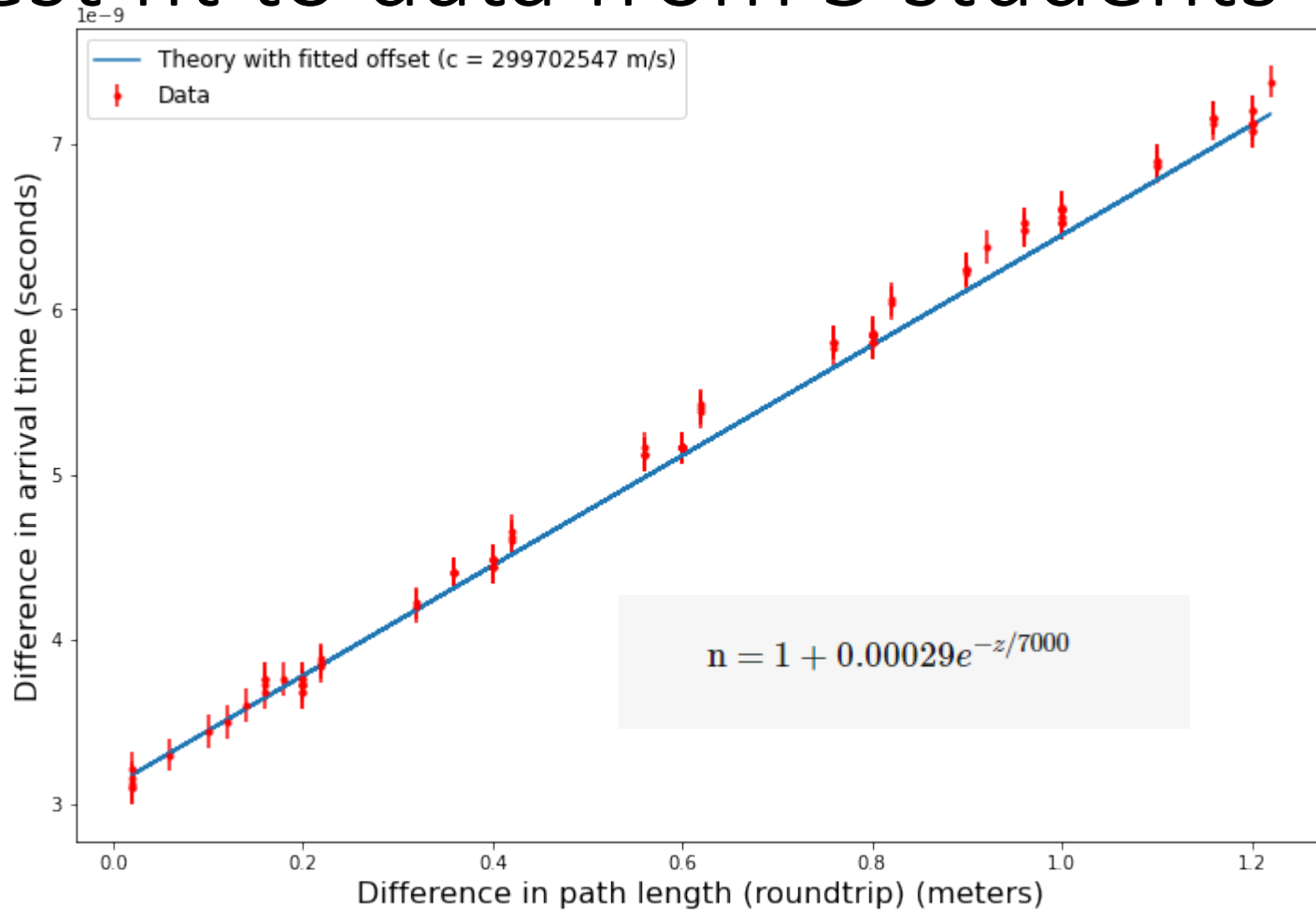


$R^2 = 0.997$

The slope = $3.4470692242258648e-09$, with uncertainty $1.741025266191759e-11$
The intercept = $3.111703798698238e-09$, with uncertainty $1.298000077865901e-11$
The speed of light = 290101513.76480633

Speed of Light Data – Our Class

Best fit to data from 5 students



Can we reject the hypothesis that the speed of light is 299,700,000 m/s in our lab?

Examples of student notebooks

$$v_0 = 0 \quad \Delta t_0 = 2.60 \text{ ns}$$

Trial 1

Δx	$\Delta(\Delta t)$
10.0 cm	0.84 ns
20.0 cm	1.68 ns
30.0 cm	2.36 ns
40.0 cm	3.08 ns
50.0 cm	3.8 ns
60.0 cm	4.44 ns

T4's was actually
1.60 and was
a typo

$$v_0 = 0 \quad \Delta t_0 = 2.64 \text{ ns}$$

Trial 2

Δx	$\Delta(\Delta t)$
10.0 cm	0.76 ns
20.0 cm	1.56 ns
30.0 cm	2.40 ns
40.0 cm	3.08 ns
50.0 cm	3.68 ns
60.0 cm	4.40 ns



$$v_0 = 0 \quad \Delta t_0 = 2.60 \text{ ns}$$

Trial 3

Δx	$\Delta(\Delta t)$
10.0 cm	0.84 ns
20.0 cm	1.68 ns
30.0 cm	2.44 ns
40.0 cm	3.16 ns
50.0 cm	3.72 ns
60.0 cm	4.44 ns

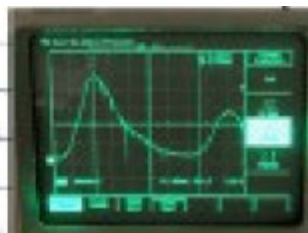
$$v_0 = 0, \quad \Delta t_0 = 2.60 \text{ ns}$$

Trial 4

Δx	$\Delta(\Delta t)$
10.0 cm	2.5 ns 0.84 ns
20.0 cm	1.60 ns
30.0 cm	2.40 ns
40.0 cm	3.16 ns
50.0 cm	3.76 ns
60.0 cm	4.44 ns

corner mirror distance	51.25 cm	61.25 cm	71.25 cm	81.25 cm
temporal separation	2.84 ns	3.64 ns	4.4 ns	5.32 ns
	$\frac{91.25 \text{ cm}}{6.06 \text{ ns}}$			

For x the position of the corner-cube reflector and t the pulse delay:							
The added distance to the light beam is roughly $2(x+5)$ cm with 5cm the distance from the beam splitter to the mirror in the setup							
Data							
		Trial 1	Trial 2	Trial 3	Trial 4	Mean	
x (m)	$2x+0.1$ (m)	t (ns)	t (ns)	t (ns)	t (ns)	t (ns)	
0.58	1.26	7.16	7.12	7.16	7.16	7.15	
0.48	1.06	6.48	6.52	6.52	6.48	6.5	
0.38	0.86	5.8	5.76	5.8	5.8	5.79	
0.28	0.66	5.16	5.12	5.12	5.12	5.13	
0.18	0.46	4.4	4.4	4.4	4.4	4.4	
0.08	0.26	3.76	3.68	3.72	3.76	3.73	



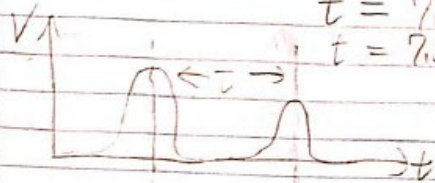
I varied the separation of the corner cube by 5 cm every measurement. Recorded below are the measurements I got at each distance.

Corner cube location	Δt trial 1	Δt trial 2	Δt trial 3	Δt trial 4
55 cm	6.90 ns ± 0.02 ns	6.86 ns ± 0.02 ns	6.90 ns ± 0.02 ns	6.88 ns ± 0.02 ns
50 cm	6.62 ns ± 0.02 ns	6.60 ns ± 0.02 ns	6.62 ns ± 0.02 ns	6.60 ns ± 0.02 ns
45 cm	6.24 ns ± 0.02 ns	6.24 ns ± 0.02 ns	6.24 ns ± 0.02 ns	6.22 ns ± 0.02 ns
40 cm	5.80 ns ± 0.02 ns	5.86 ns ± 0.02 ns	5.86 ns ± 0.02 ns	5.84 ns ± 0.02 ns

Position 1: $x = 61.0 \pm 0.1 \text{ cm}$

$t = 7.12 \pm 0.01 \text{ ns (Air)}$

$t = 7.56 \pm 0.01 \text{ ns (rod)}$



Position 2: $x = 52.8 \pm 0.1 \text{ cm}$

Set 1: $t = 6.60 \pm 0.01 \text{ ns (Air)}$ $t = 6.92 \pm 0.01 \text{ ns (rod)}$

Position 3: $x = 49.0 \pm 0.1 \text{ cm}$

$t = 6.12 \pm 0.01 \text{ ns (Air)}$ $t = 6.52 \pm 0.01 \text{ ns (rod)}$

Position 4: $x = 37.0 \pm 0.1 \text{ cm}$

$t = 5.52 \pm 0.01 \text{ ns (Air)}$ $t = 5.96 \pm 0.01 \text{ ns (rod)}$

Position 5: $x = 30.0 \pm 0.1 \text{ cm}$

$t = 5.12 \pm 0.01 \text{ ns (Air)}$ $t = 5.52 \pm 0.01 \text{ ns (rod)}$

Position 1: $x = 55.0 \pm 0.1 \text{ cm}$

$t = 6.80 \pm 0.01 \text{ ns (Air)}$ $t = 7.24 \pm 0.01 \text{ ns (rod)}$

Position 2: $x = 50.0 \pm 0.1 \text{ cm}$

$t = 6.48 \pm 0.01 \text{ ns (Air)}$ $t = 6.89 \pm 0.01 \text{ ns (rod)}$

Set 2

Position 3: $x = 45.0 \pm 0.1 \text{ cm}$

$t = 6.20 \pm 0.01 \text{ ns (Air)}$ $t = 6.52 \pm 0.01 \text{ ns (rod)}$

Position 4: $x = 40.0 \pm 0.1 \text{ cm}$

$t = 5.80 \pm 0.01 \text{ ns (Air)}$ $t = 6.16 \pm 0.01 \text{ ns (rod)}$

Position 5: $x = 35.0 \pm 0.1 \text{ cm}$

$t = 5.52 \pm 0.01 \text{ ns (Air)}$ $t = 5.80 \pm 0.01 \text{ ns (rod)}$