

Astro 426/526

Fall 2019

Prof. Darcy Barron

Lecture 23: Sub-millimeter and Millimeter-Wave Detection

Project Update

- Research-style paper summarizing your methods and results is due **Monday, November 25 at 4pm**
 - Group submission *on Learn* (as pdf or word document)
 - I will make sure to allow unlimited submission attempts
- Final deadline for revisions to “Results” section: Tuesday, December 3 at 12pm (*send via email*)
- Analysis code (*submit via email*) and project contribution statement (*submit on Learn*) due at start of class, Wednesday, December 4, 2019
- All three groups will give in-class presentations on Wednesday, December 4 (25 minutes per group)
- Project guidelines and grading posted on Learn

Syllabus update

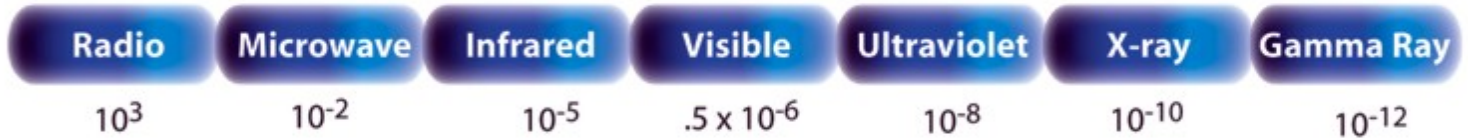
- No class Wednesday November 27 (Wednesday before Thanksgiving break)
- 6 regular classes left, covering material from *Measuring the Universe*
 - Today, Wednesday Nov 13 (spectroscopy: Chapter 6)
 - Mon Nov 18, Wed Nov 20 (mm, sub-mm: Chapter 7)
 - Monday November 25 (TBD)
 - Monday December 2 (TBD)
- Take-home final will be handed out at final class (Wednesday Dec 4), due by 12pm Wed Dec 11

THE ELECTROMAGNETIC SPECTRUM

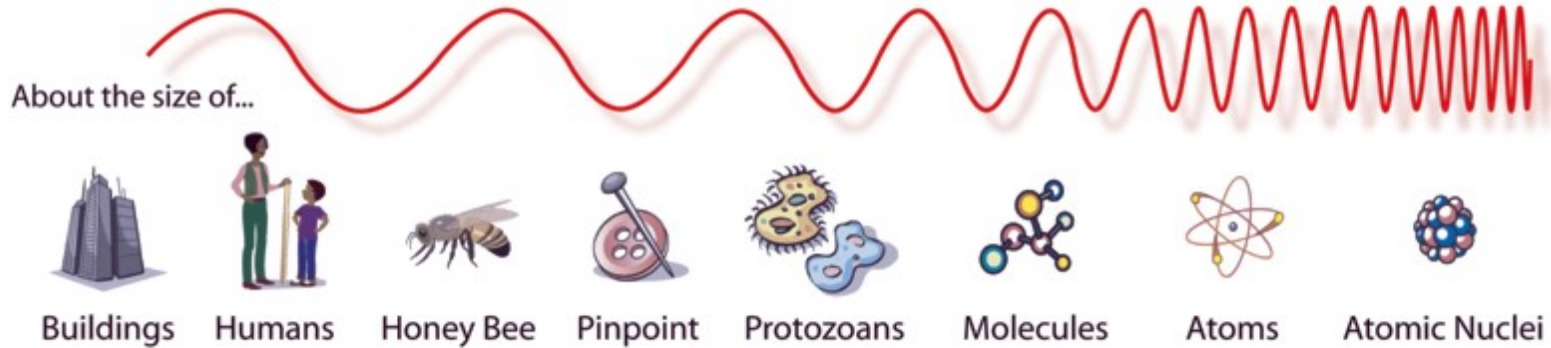
Penetrates Earth Atmosphere?



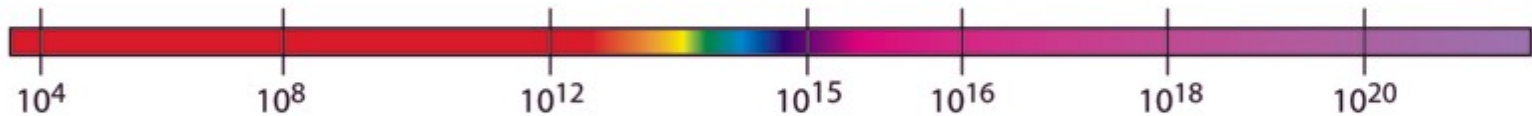
Wavelength (meters)



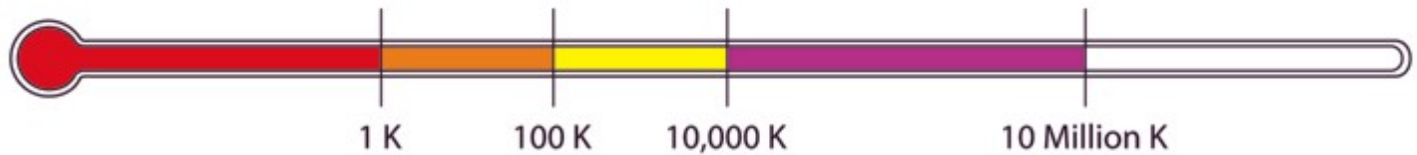
About the size of...



Frequency (Hz)



Temperature of bodies emitting the wavelength (K)



Source: NASA

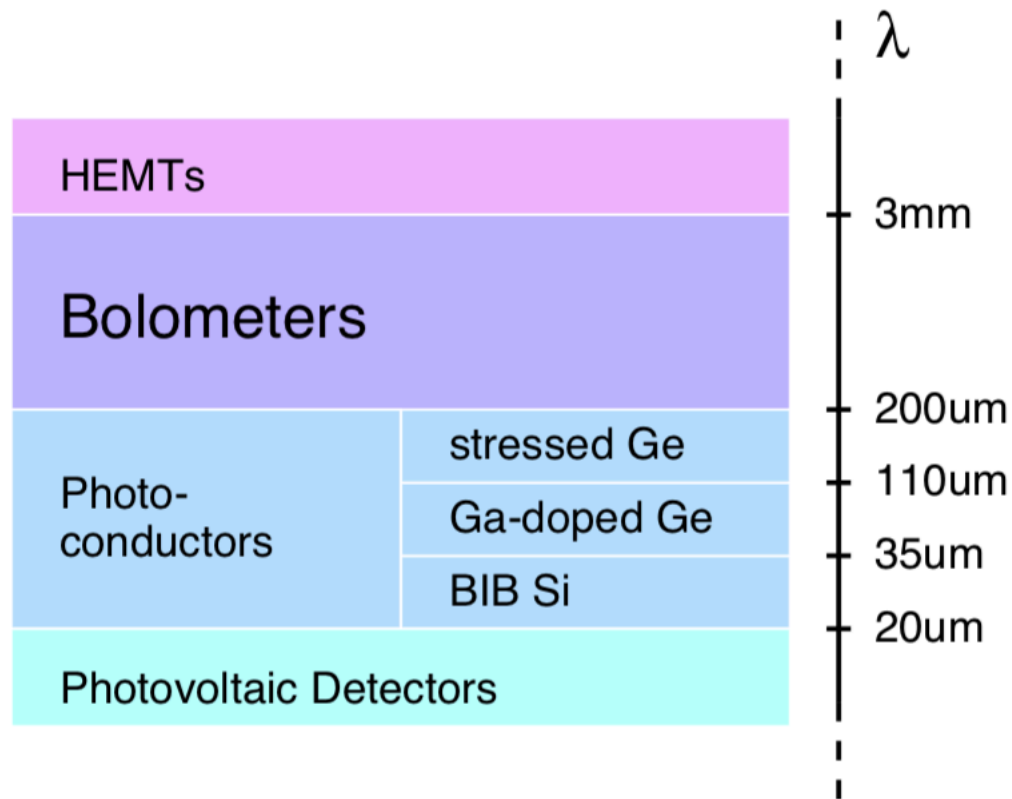


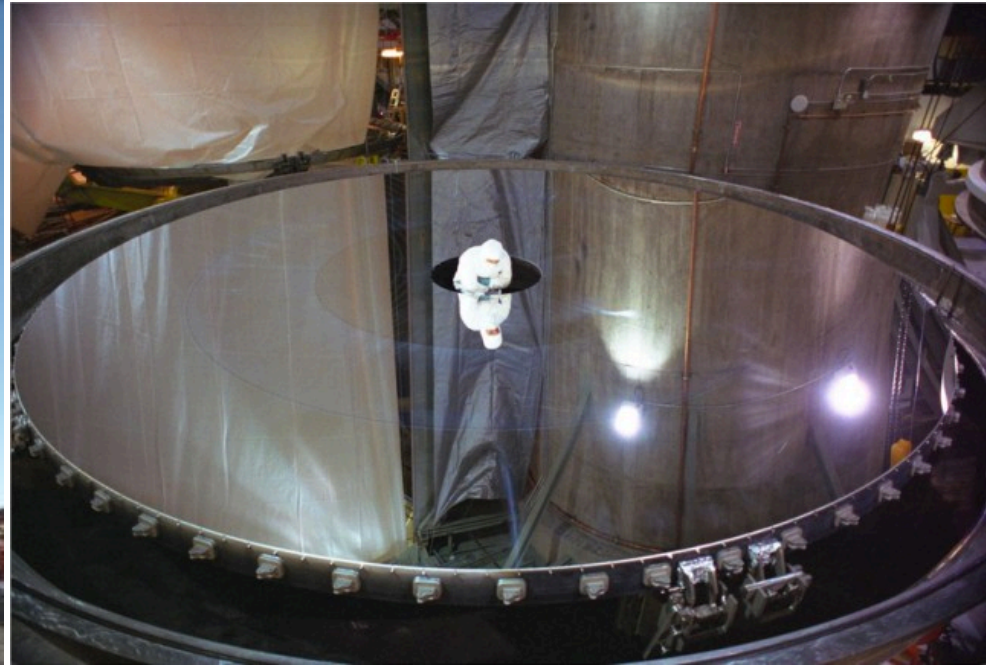
Figure 1.1: An overview of different broad band receiver technologies and the range of wavelengths in which they are predominantly used.

Figure from Gildemeister PhD Thesis, 2000

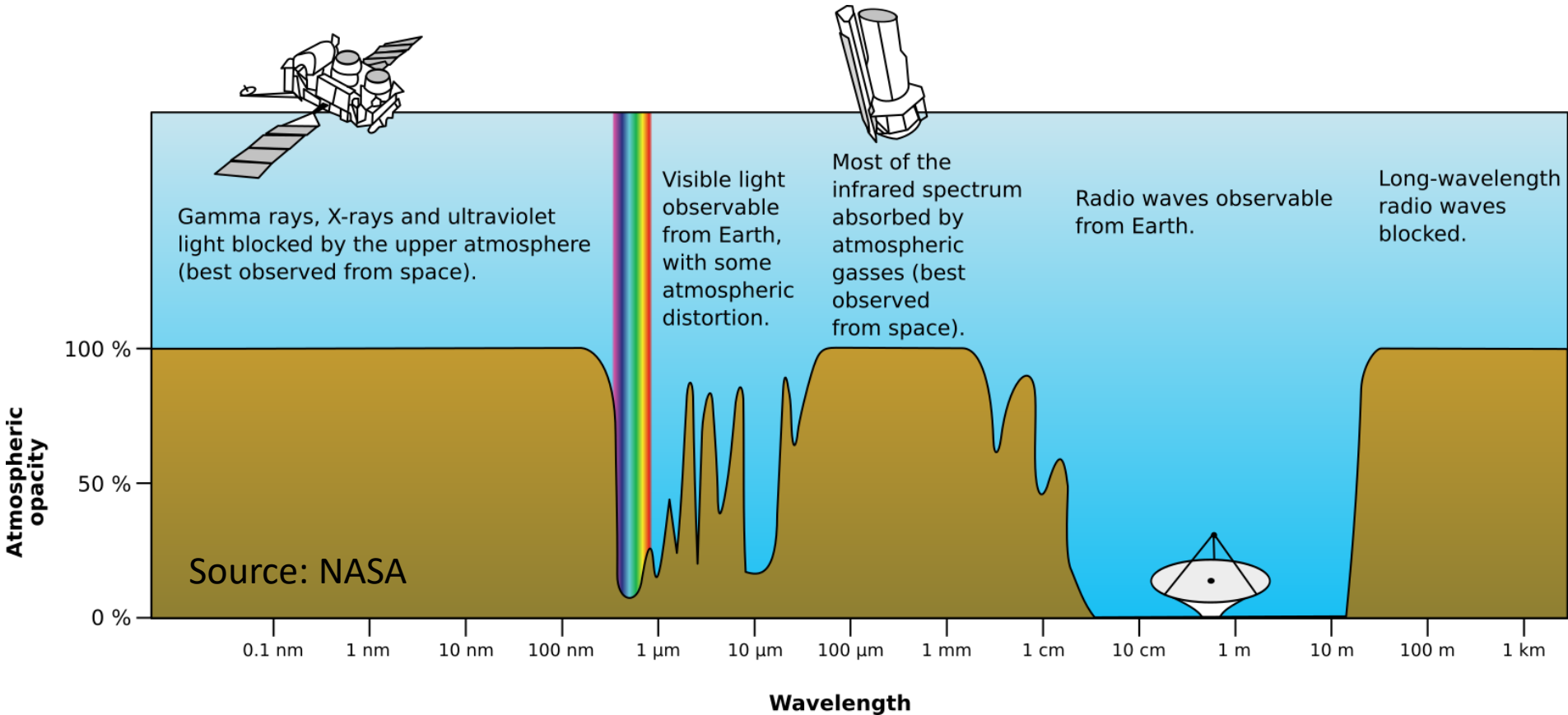
APEX telescope (sub-mm and mm)



Gemini mirror (optical)



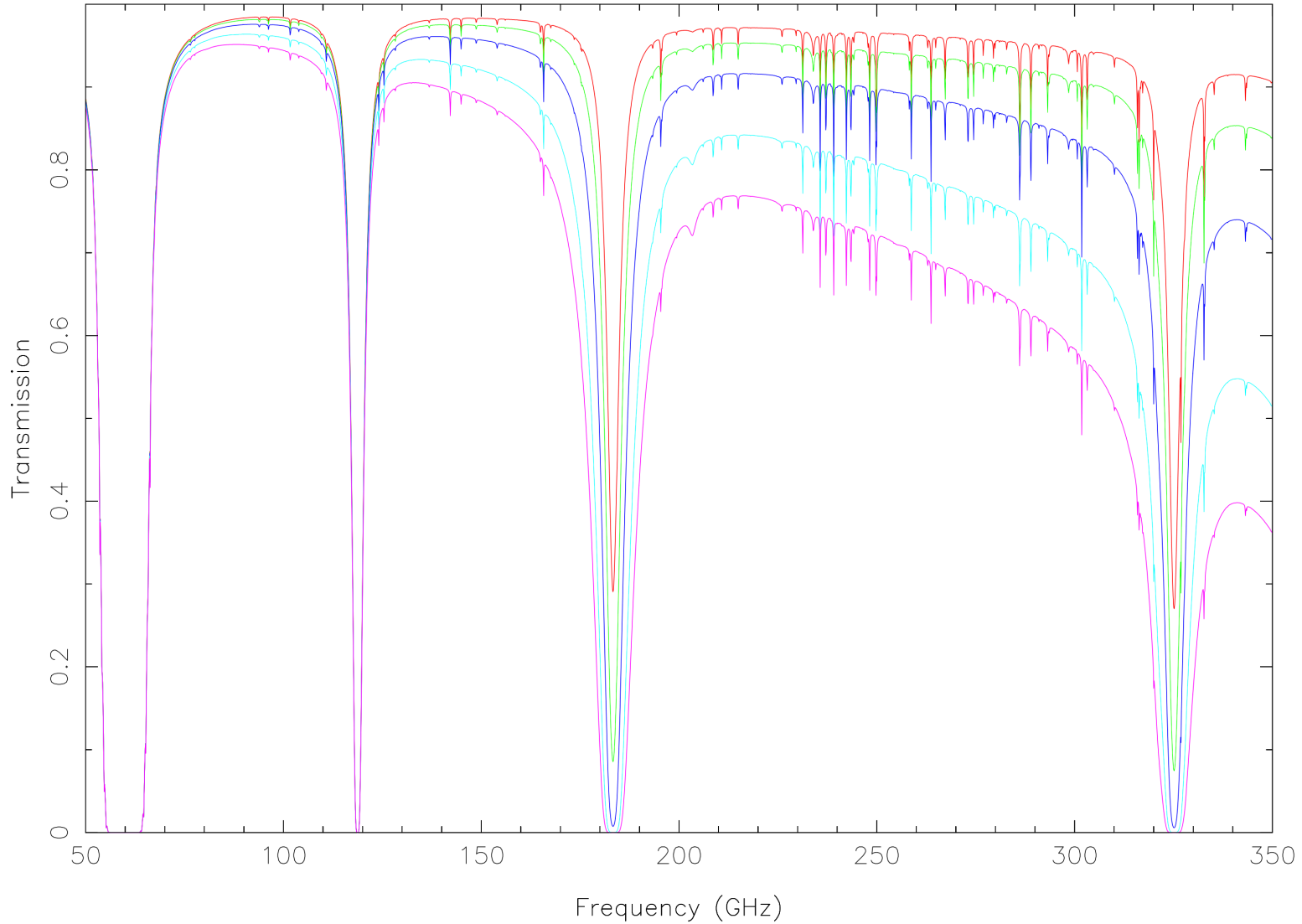
Electromagnetic spectrum and our atmosphere



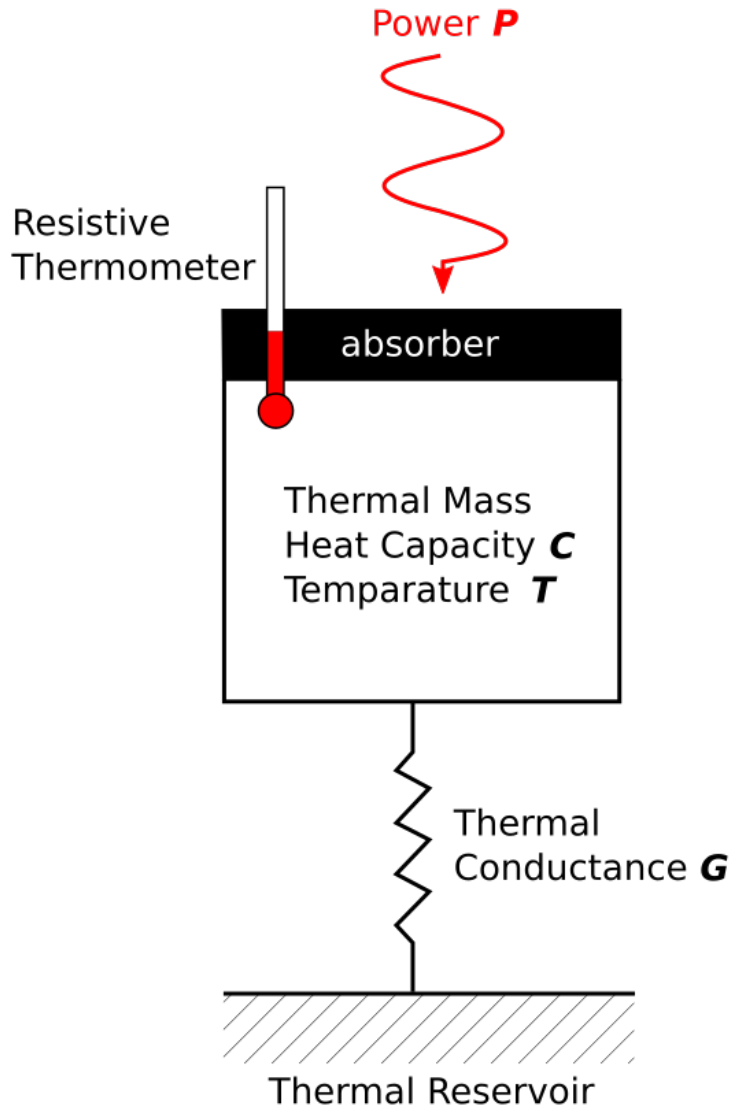
Atmospheric transmission for microwaves

APEX, Llano de Chajnantor, alt. 5109m

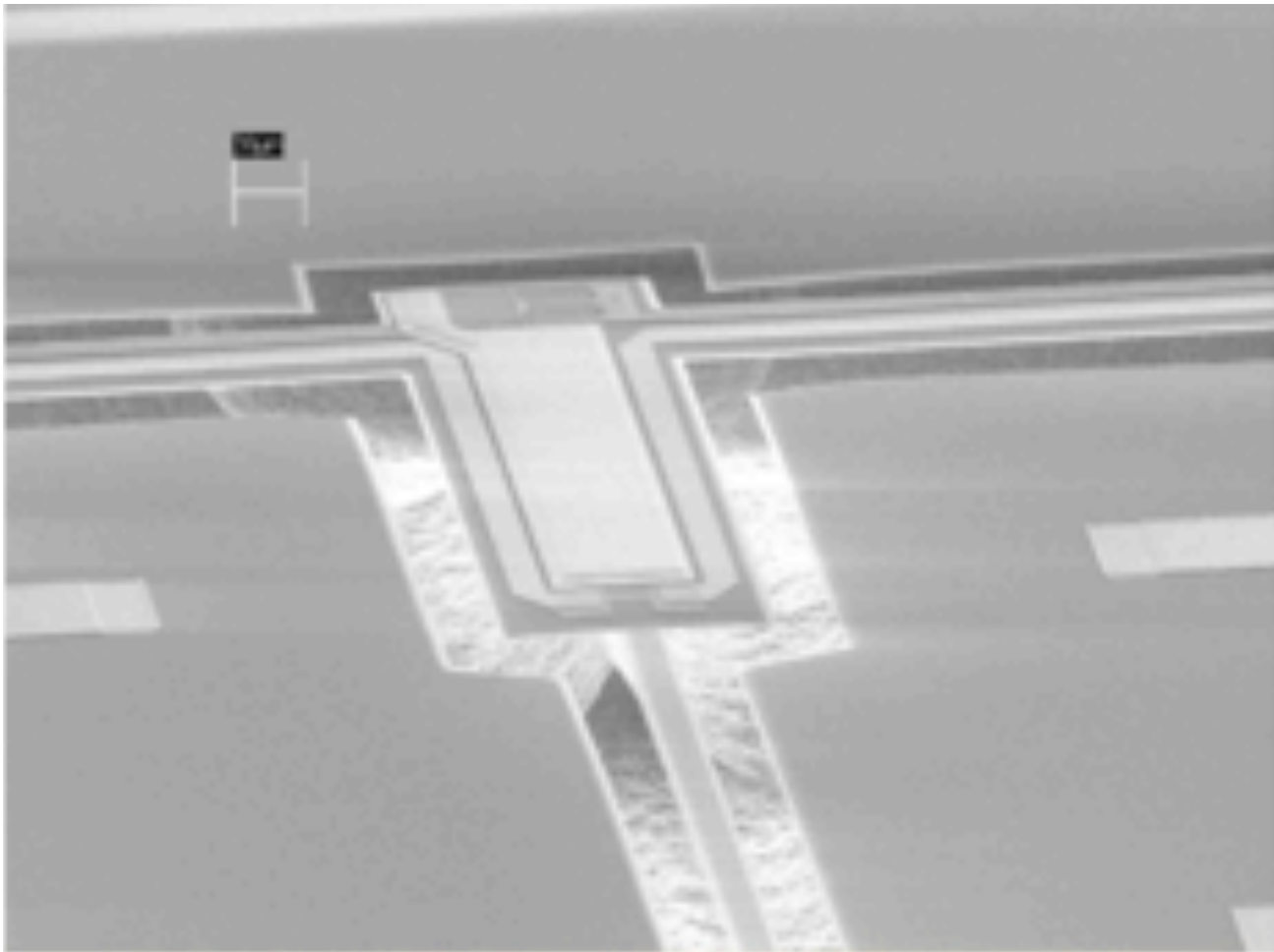
PWV=0.50 PWV=1.00 PWV=2.00 PWV=4.00 PWV=6.00

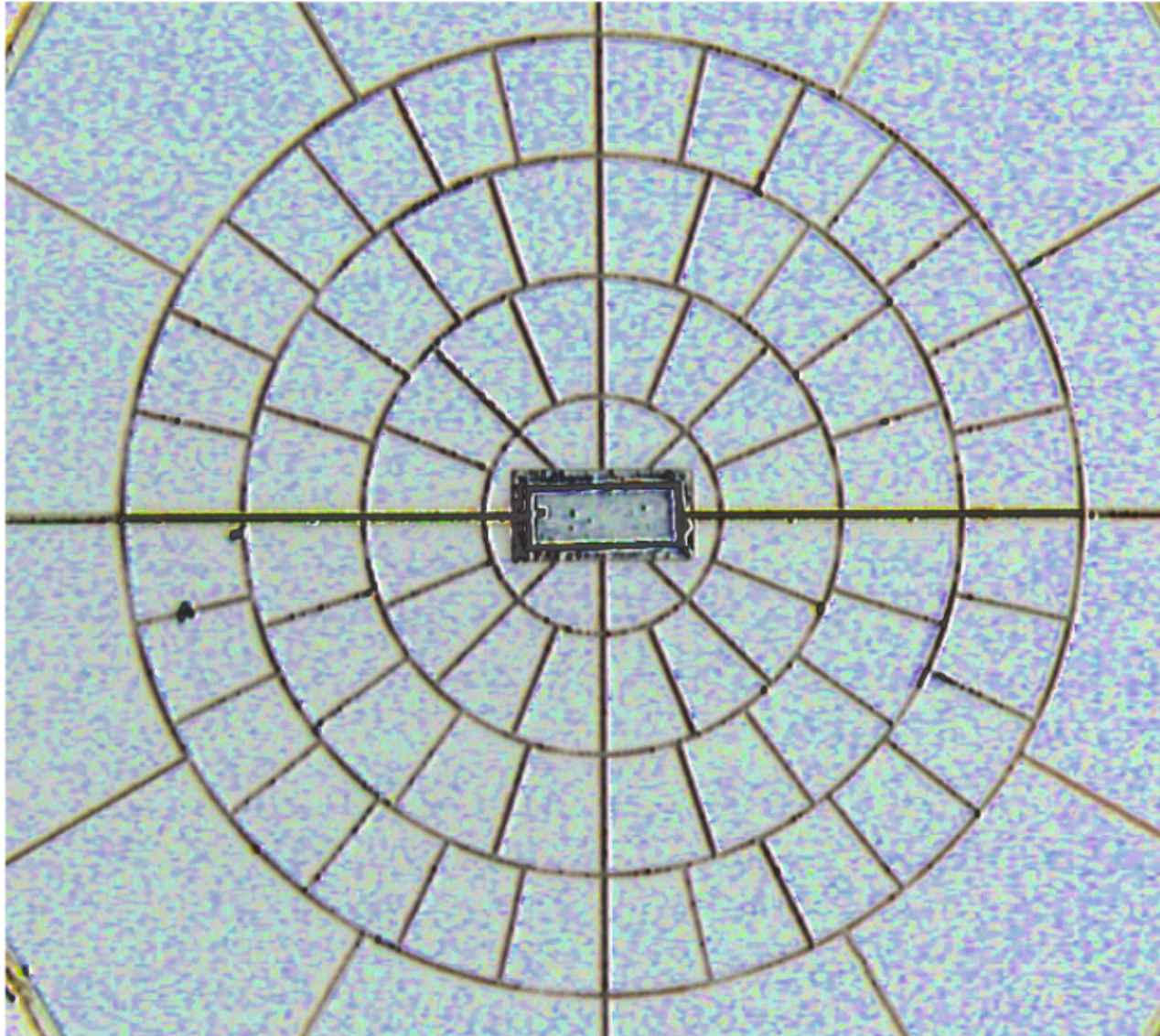


Bolometer: device to measure incident power

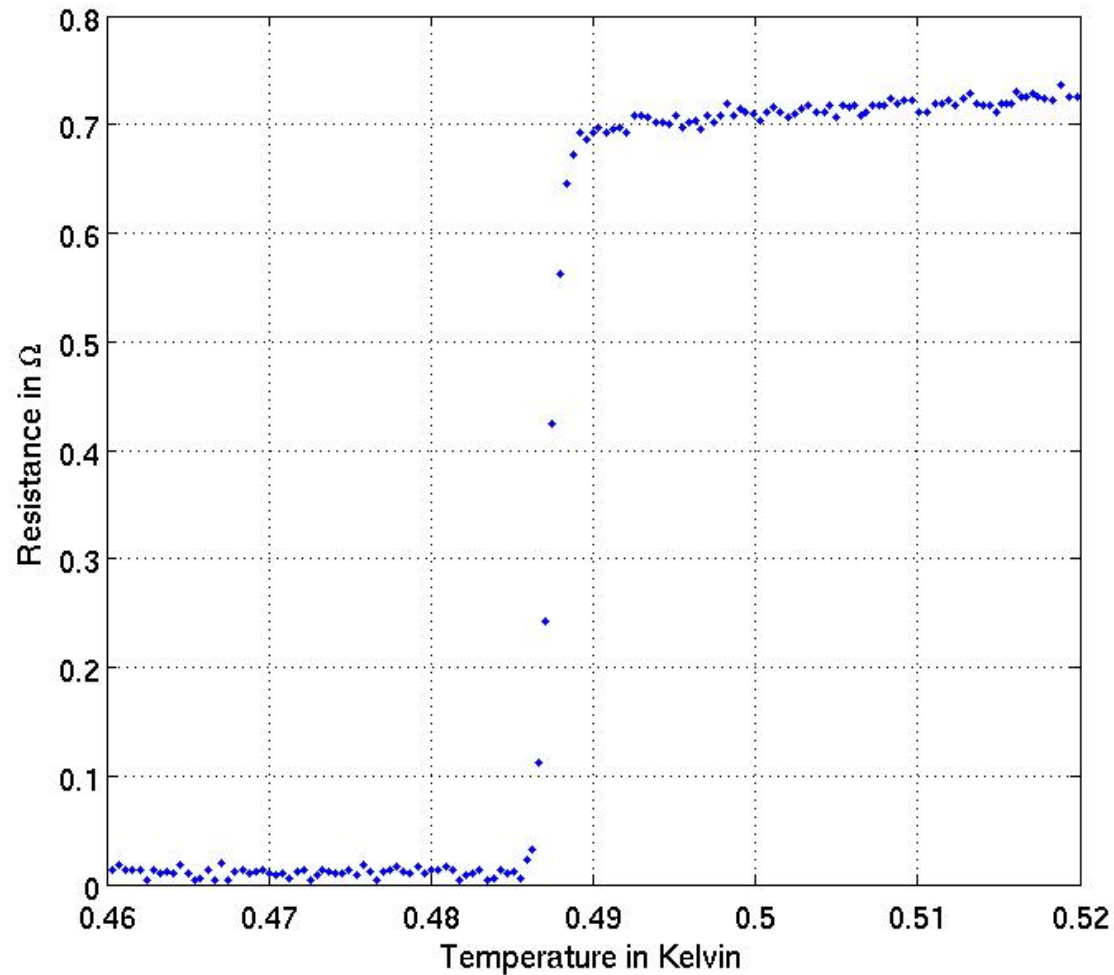


Bolometer



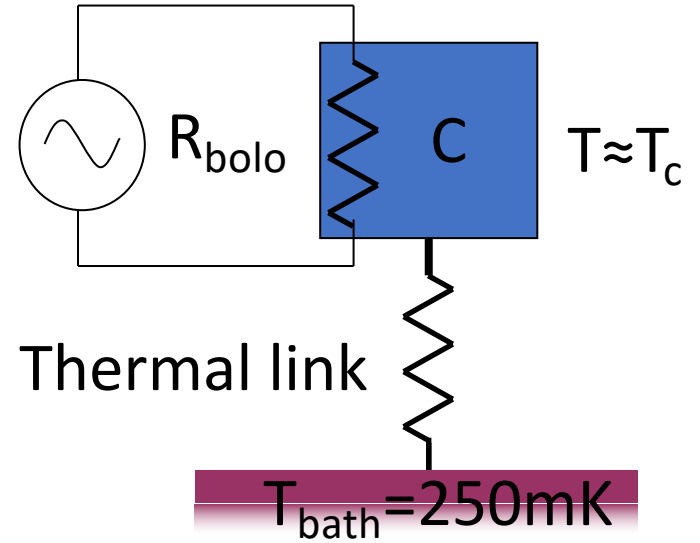


Superconducting transition (AlTi)

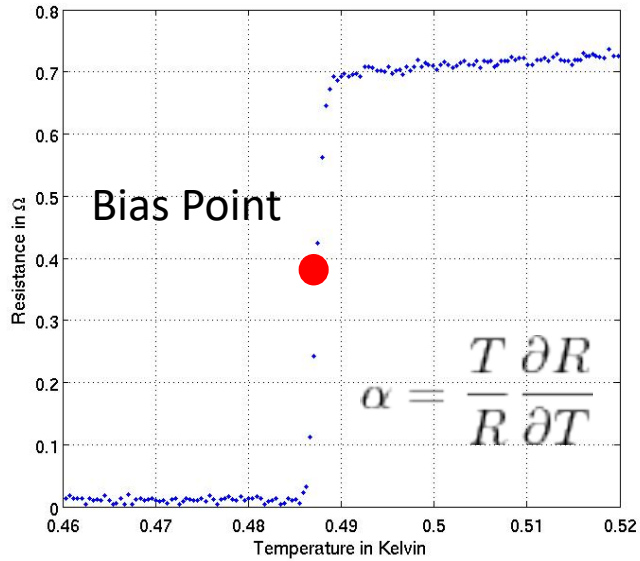


Transition Edge Sensor Bolometers

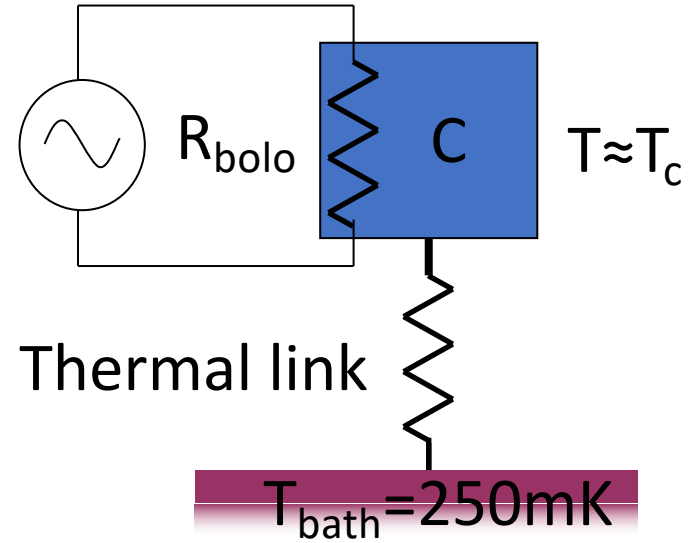
$$P_{\text{elec}} = V_{\text{bias}}^2 / R$$



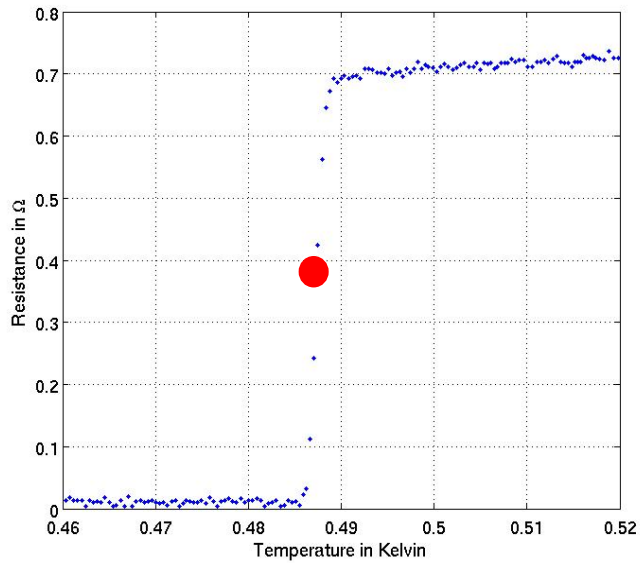
Transition Edge Sensor Bolometers



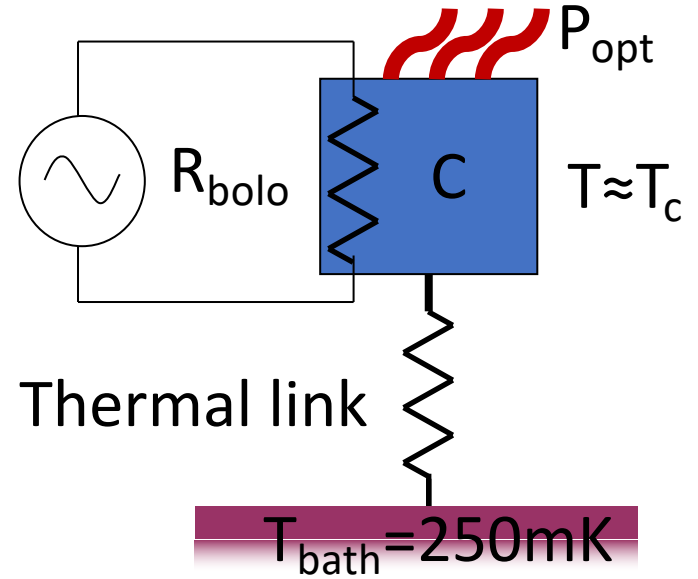
$$P_{\text{elec}} = V_{\text{bias}}^2 / R$$



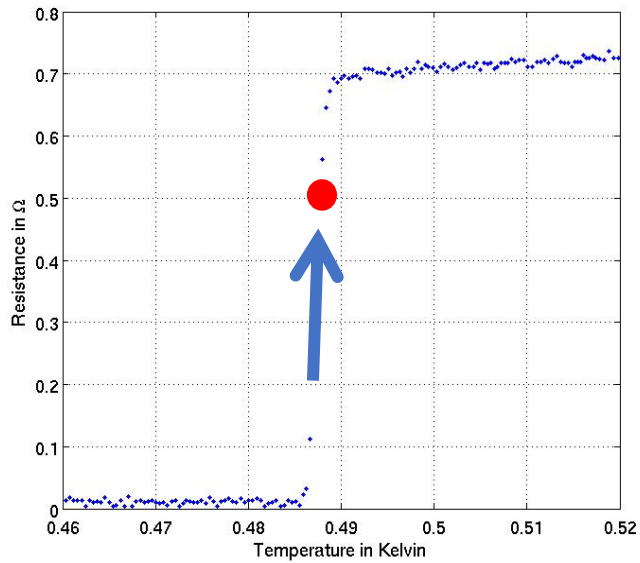
Transition Edge Sensor Bolometers



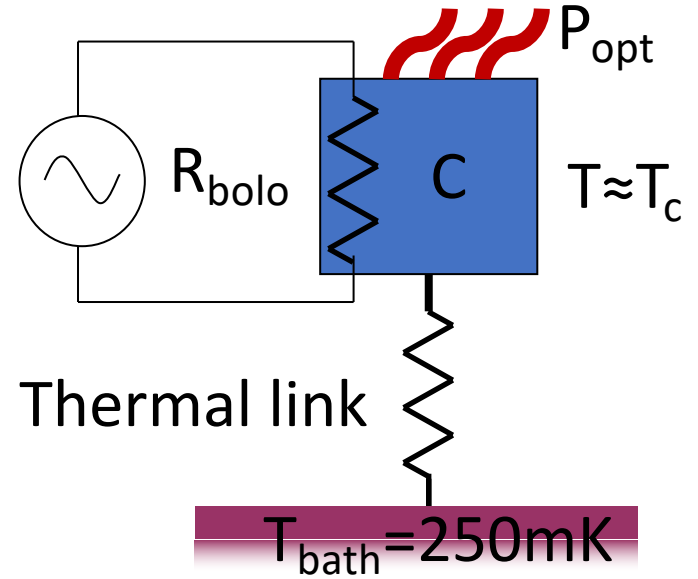
$$P_{\text{elec}} = V_{\text{bias}}^2 / R$$



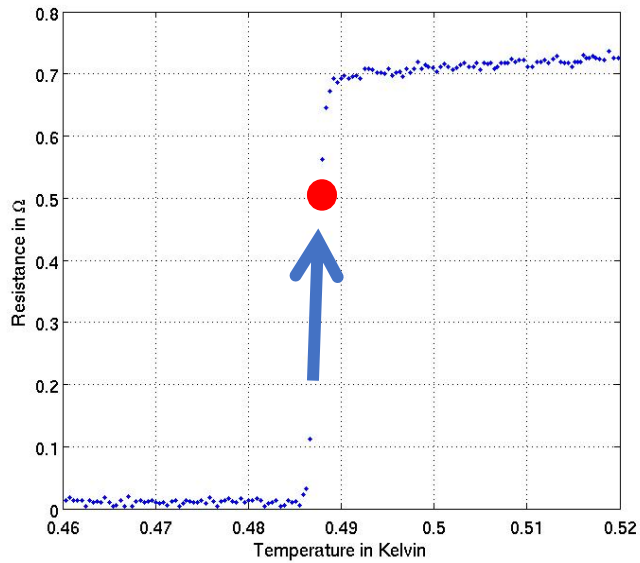
Transition Edge Sensor Bolometers



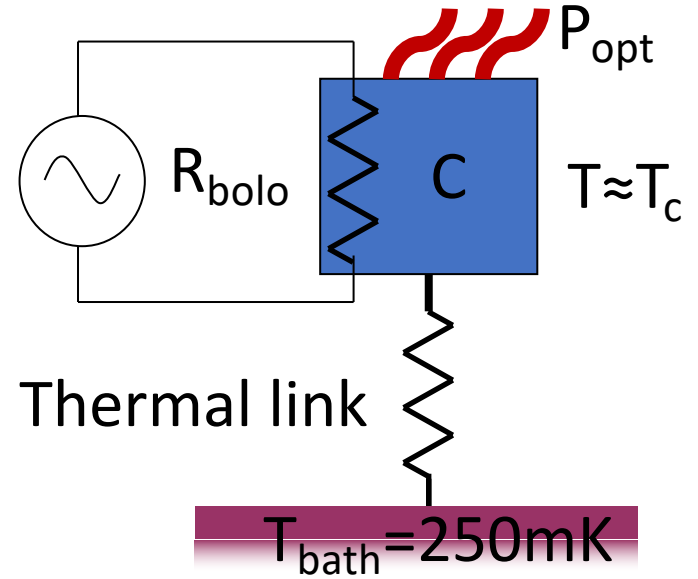
$$P_{\text{elec}} = V_{\text{bias}}^2 / R$$



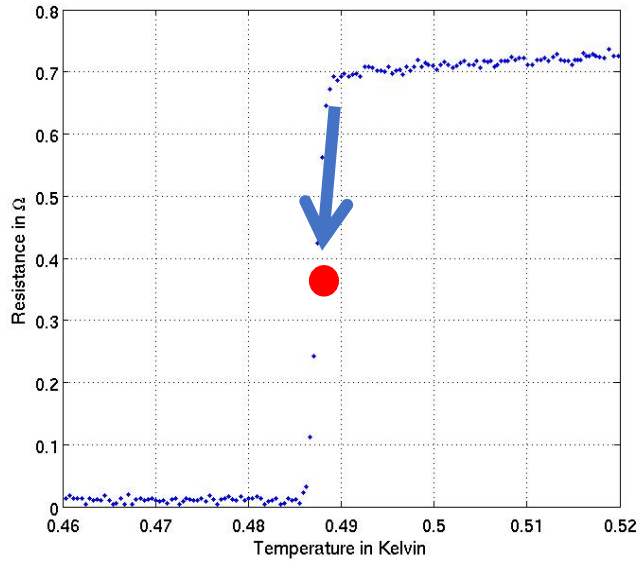
Transition Edge Sensor Bolometers



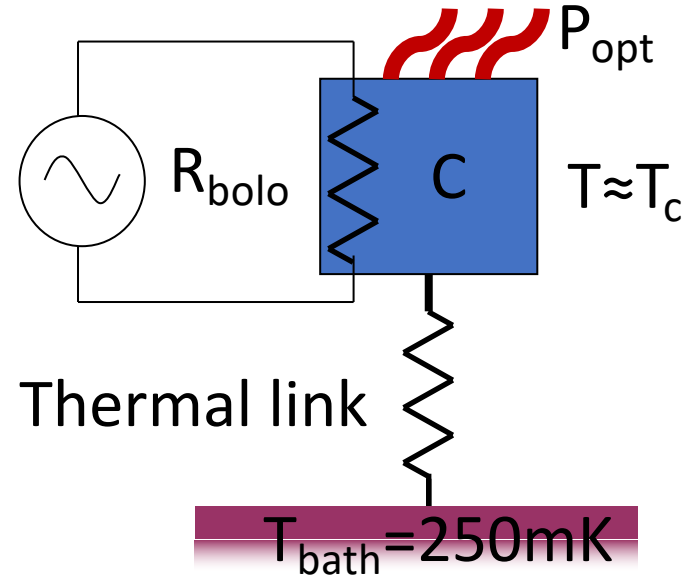
$$P_{\text{elec}} = V_{\text{bias}}^2 / R$$



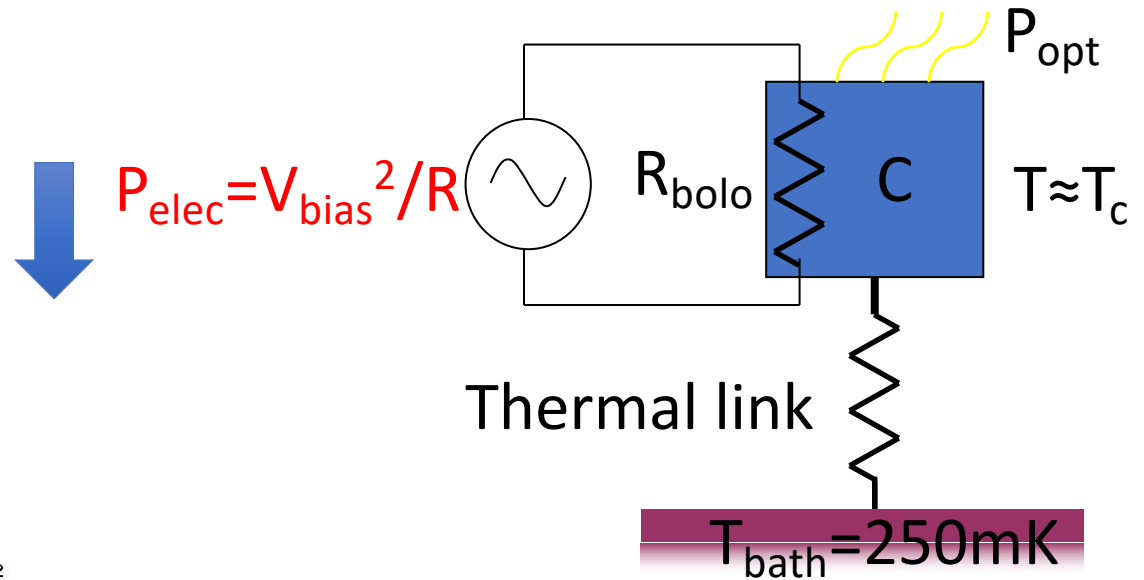
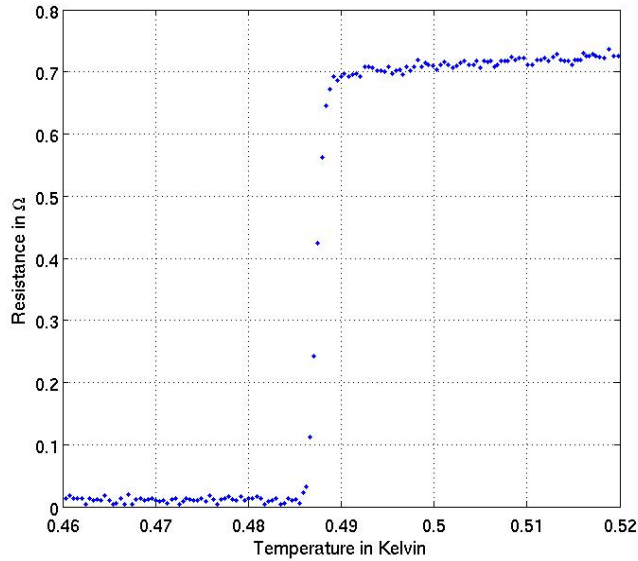
Transition Edge Sensor Bolometers



$$P_{\text{elec}} = V_{\text{bias}}^2 / R$$

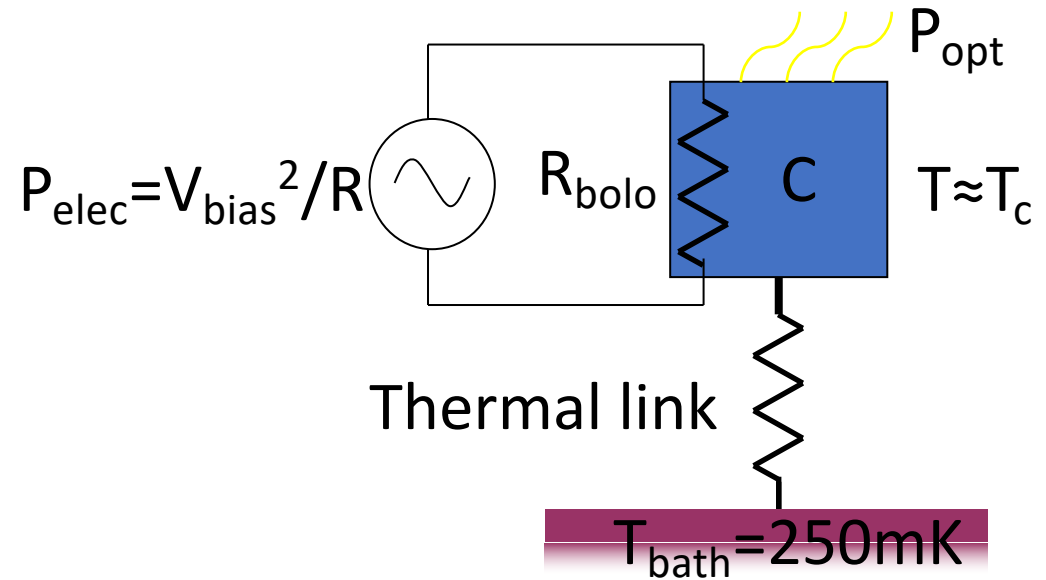
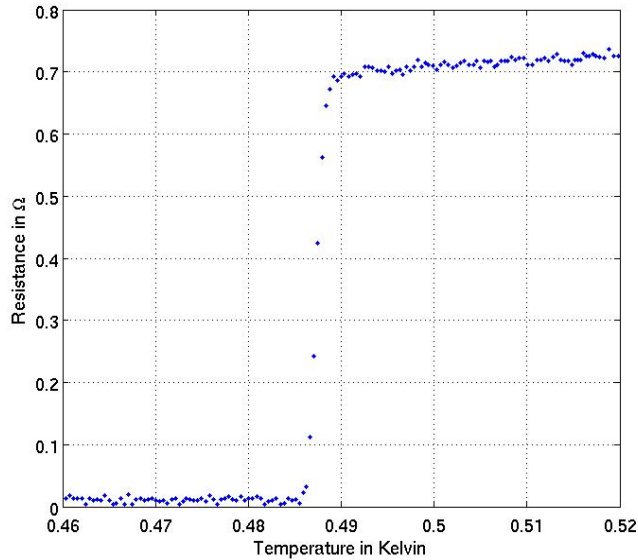


Transition Edge Sensor Bolometers



$$P_{\text{total}} = P_{\text{optical}} + P_{\text{electrical}} \approx \text{Constant}$$

Transition Edge Sensor Bolometers



$$P_{\text{total}} = P_{\text{optical}} + P_{\text{electrical}} \approx \text{Constant}$$

- Steep R vs T of superconducting transition allows sensitive detection of power on bolometer
- Strong negative feedback mechanism holds total power on bolometer constant:
Optical power modulates electrical power

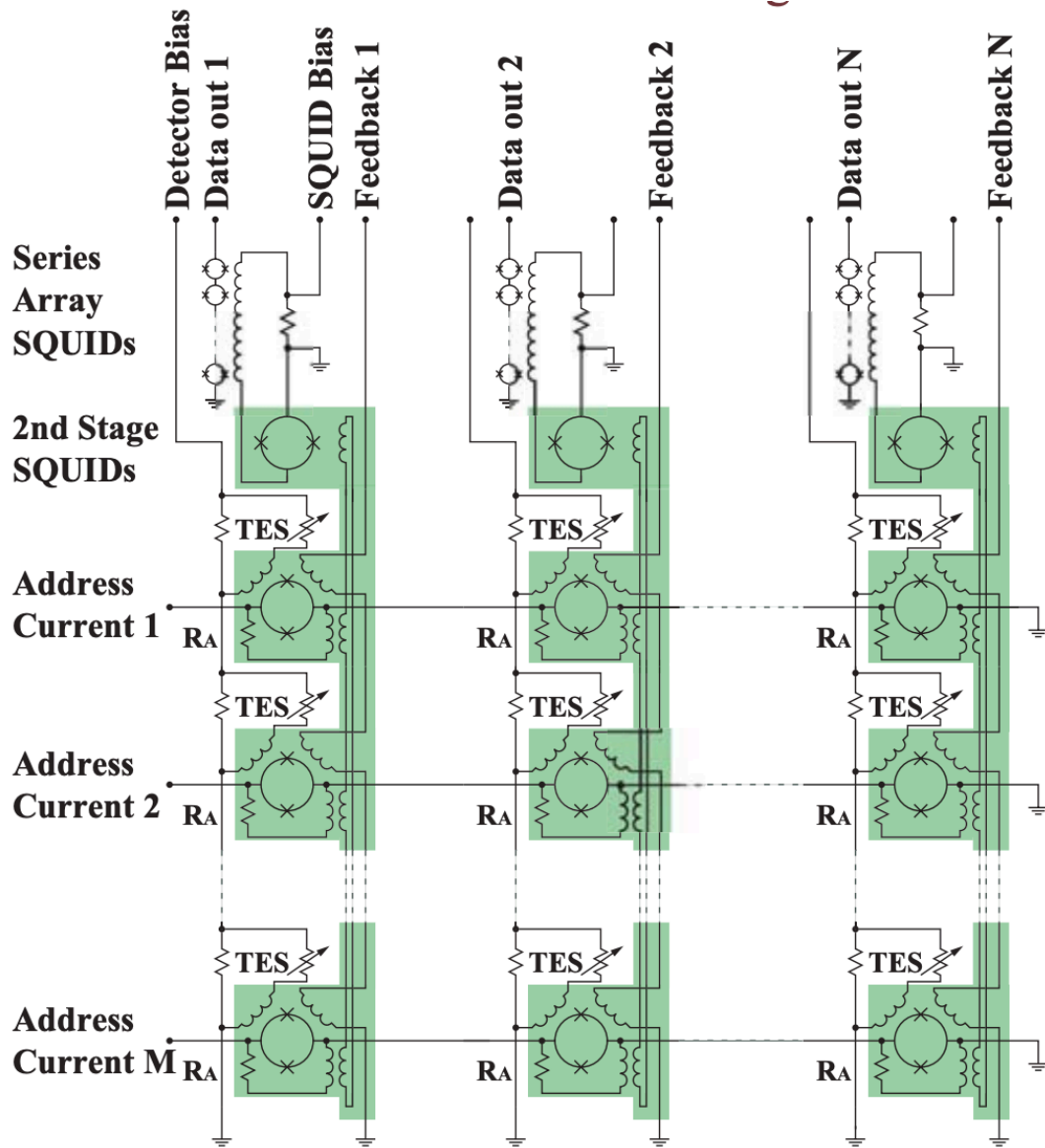
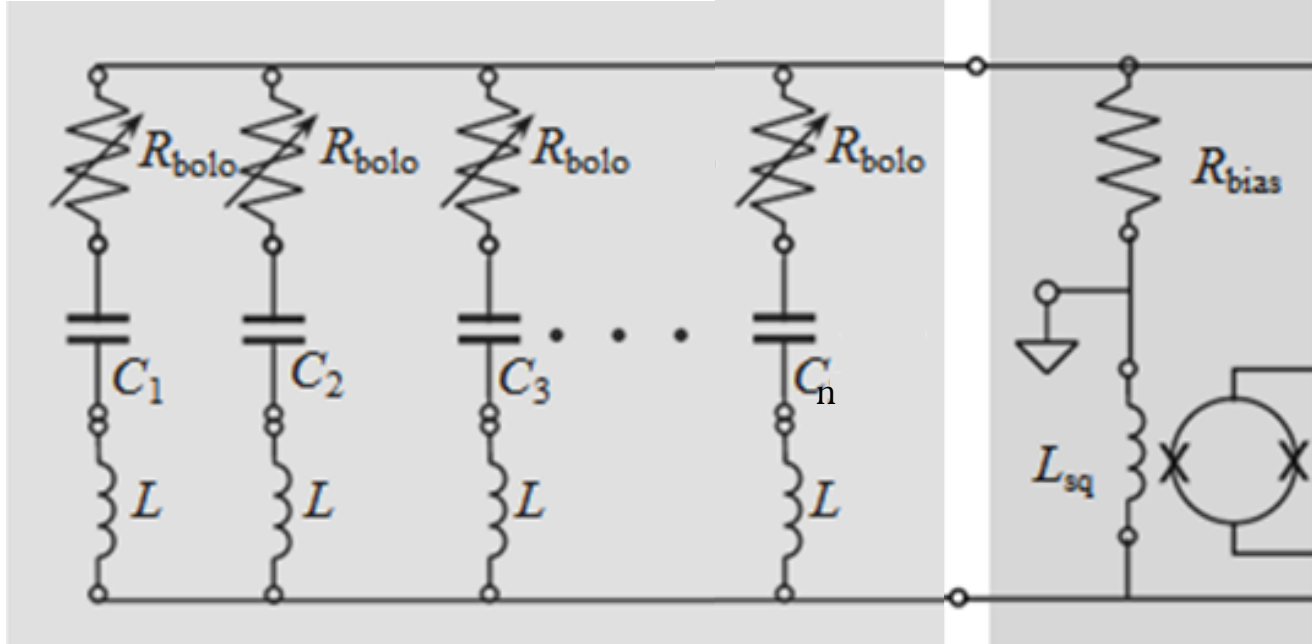
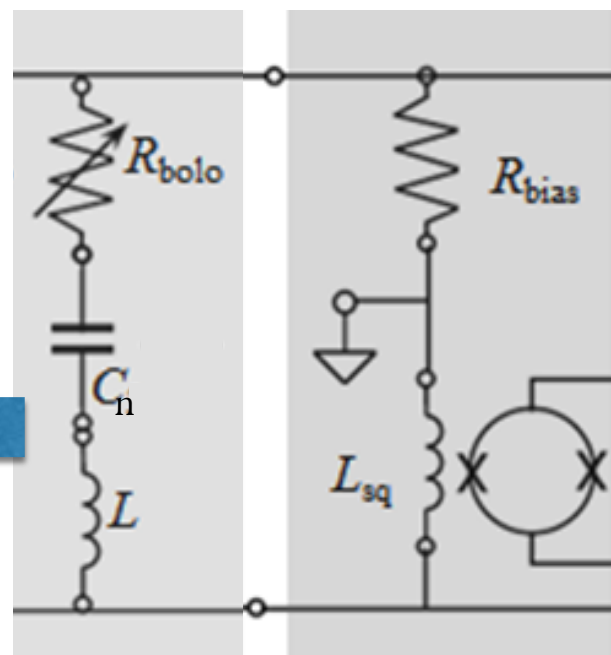
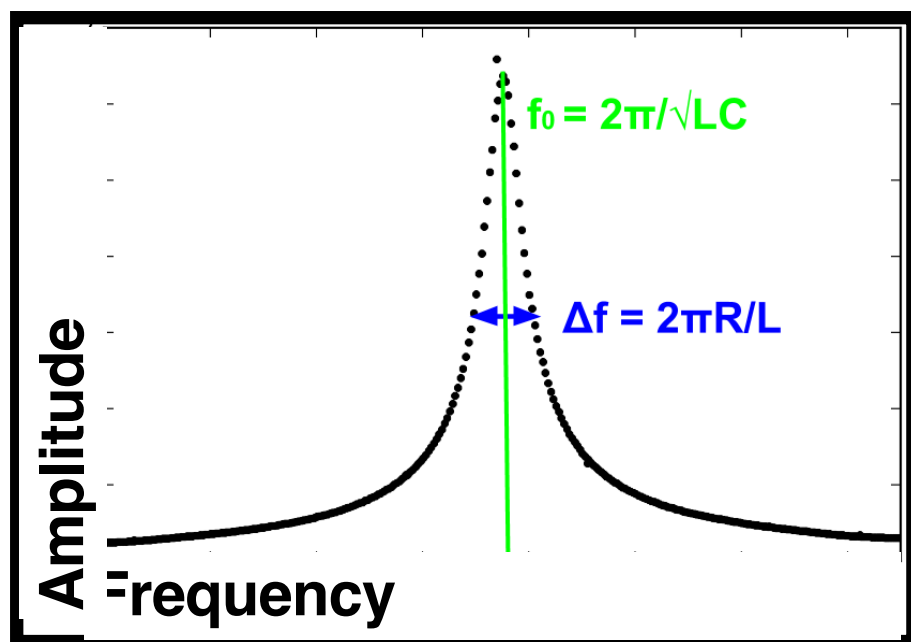


FIGURE 1. Circuit diagram for series-address SQUID MUX with $M \times N$ pixels. In this version, the first-stage SQUIDs are coupled through a common transformer to the second stage.



250 mK

4 K



250 mK

4 K